



# Latest Results from DEAP-3600



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*On Behalf of the DEAP Collaboration*  
*ICHEP 2022*  
*Bologna, Italy*



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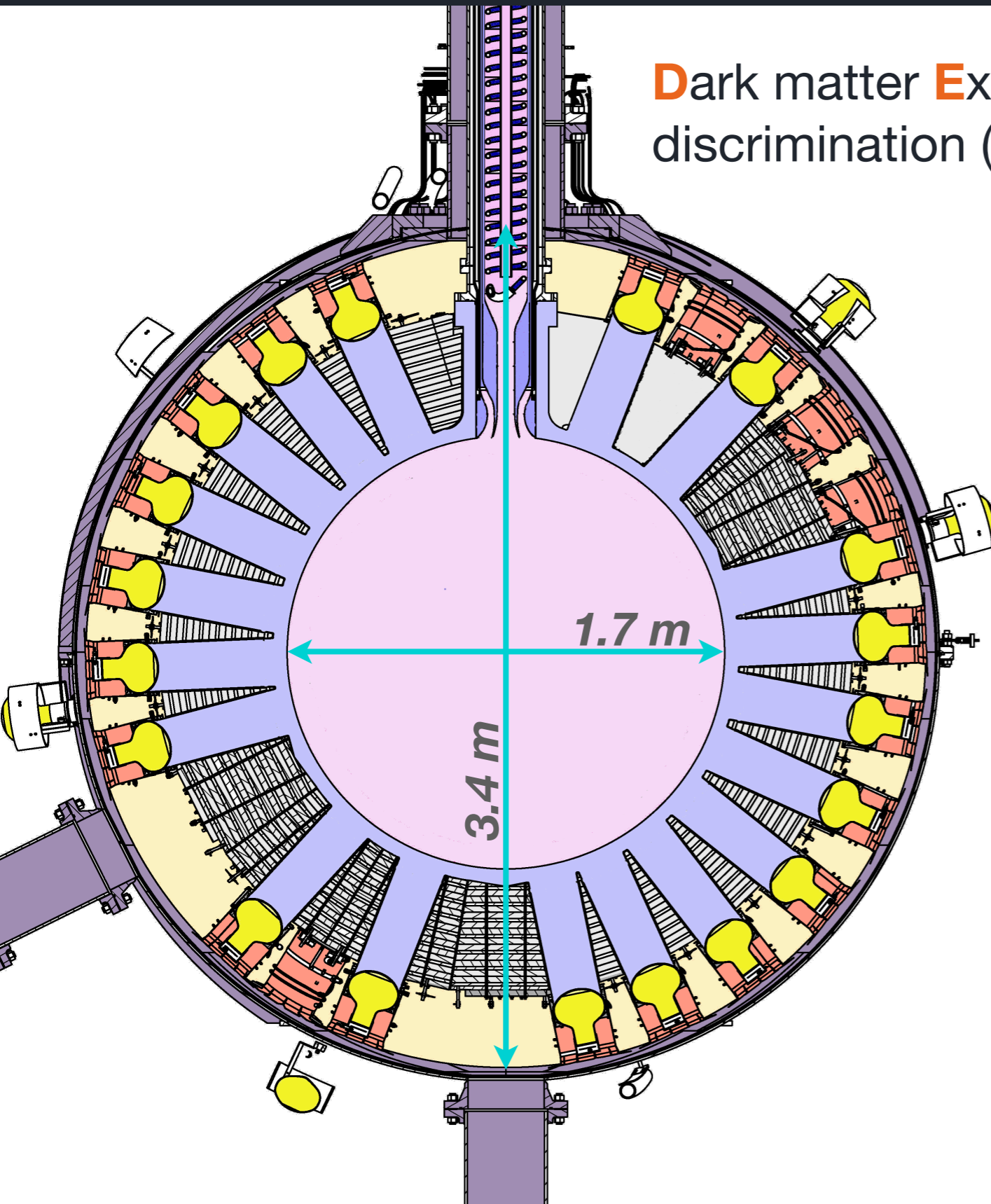


- Overview of the DEAP-3600 Experiment
- Precision Measurements
- WIMP Searches
- Beyond WIMPs
- Outlook

# Overview of DEAP-3600: The Detector



Dark matter Experiment using Argon Pulse shape discrimination (PSD)

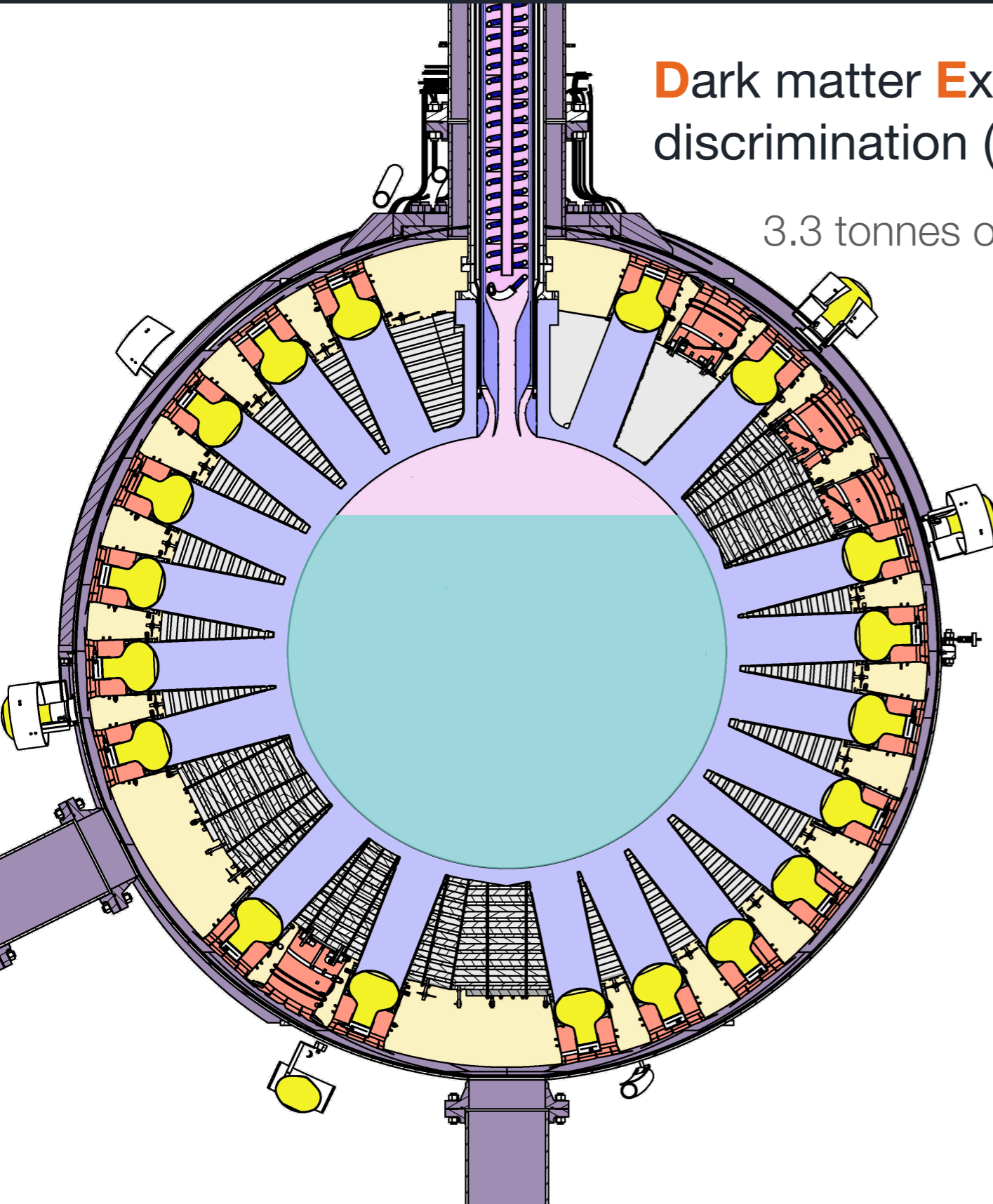


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Dark matter Experiment using Argon Pulse shape discrimination (PSD)

3.3 tonnes of Liquid Argon (LAr) as target



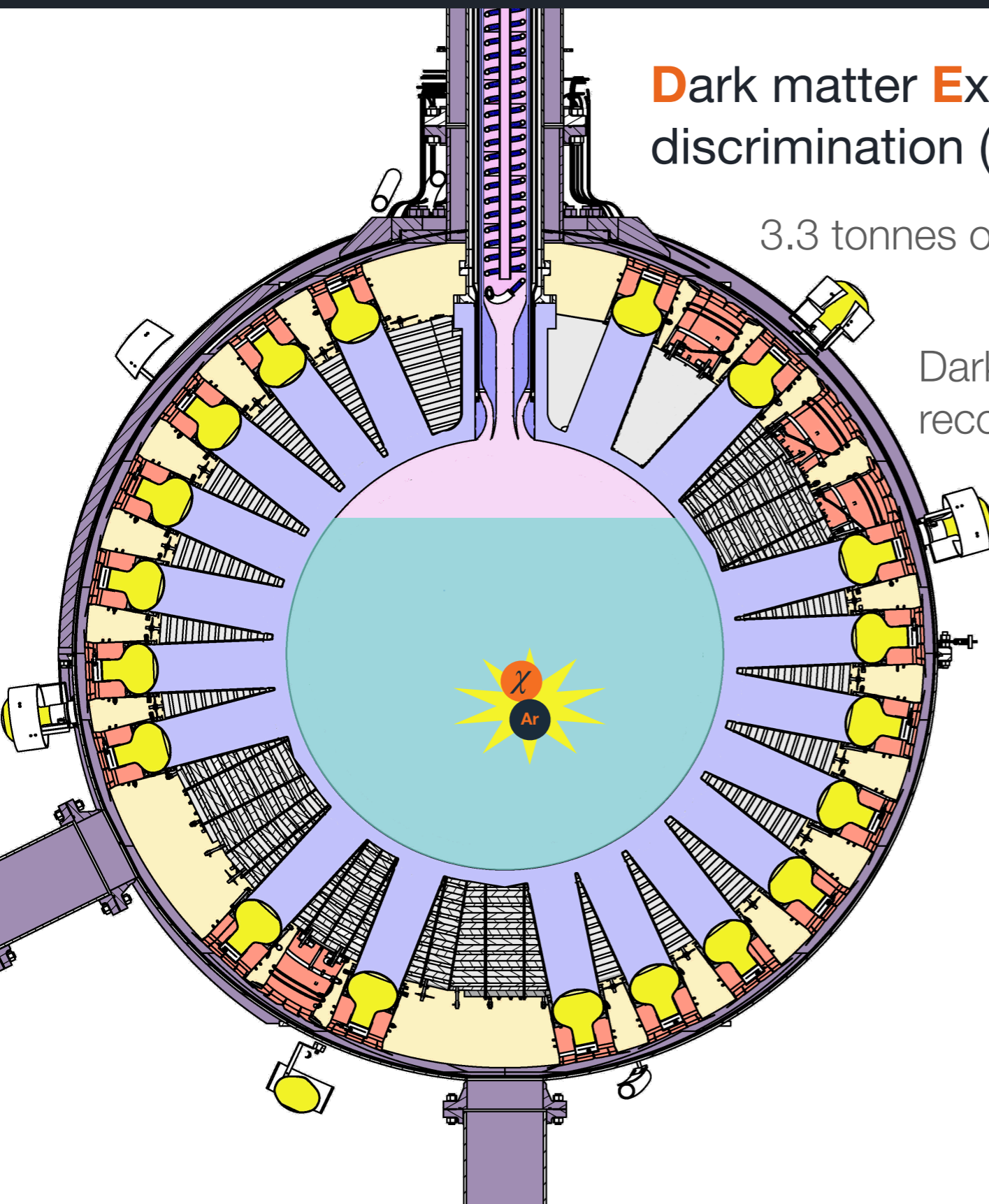
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Dark matter elastically scatters off argon nucleus, recoiling argon excites/ionizes nearby argon atoms



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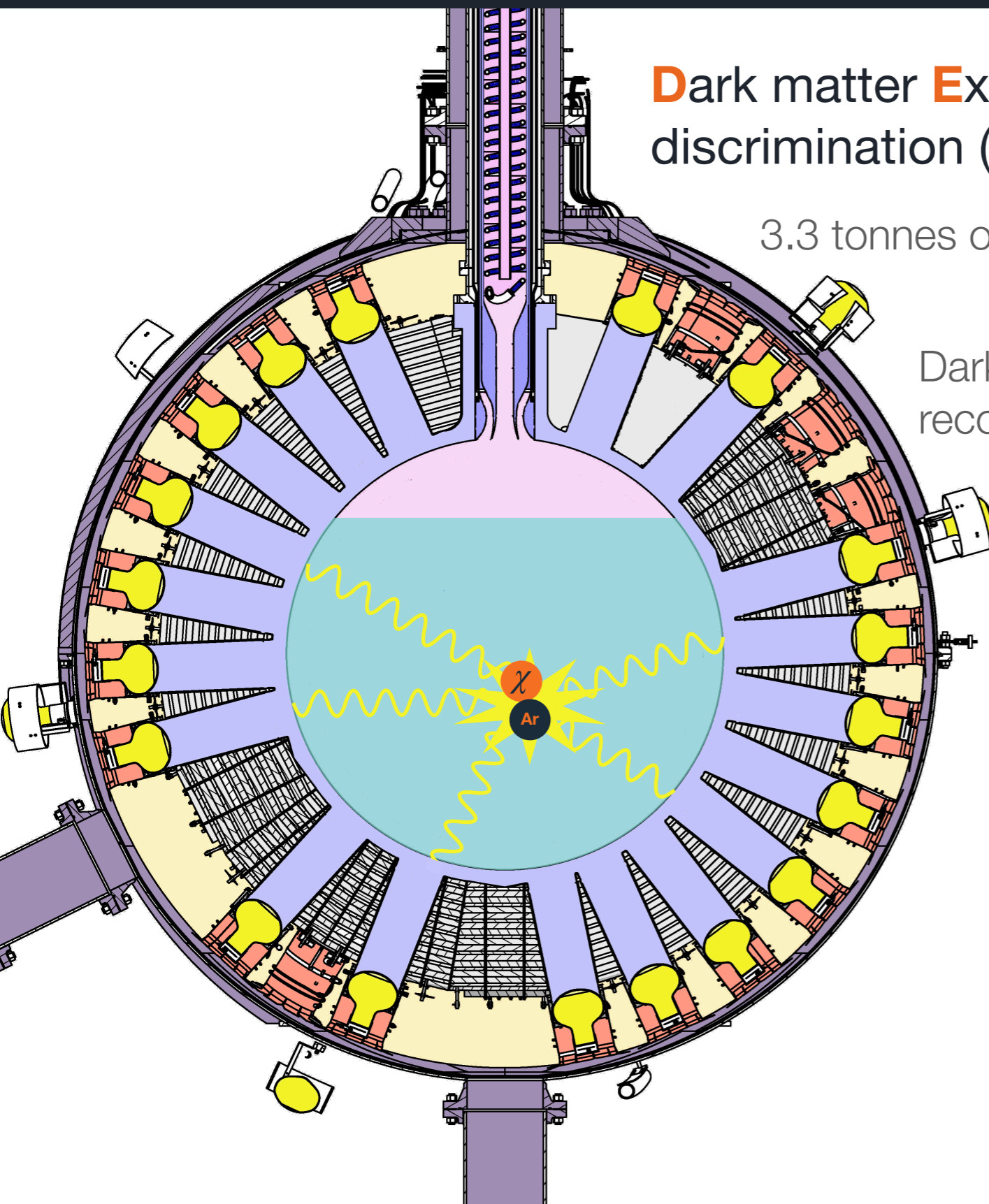


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Excited/ionized argon form excited dimer states, relax to ground state via scintillation of 128 nm photons



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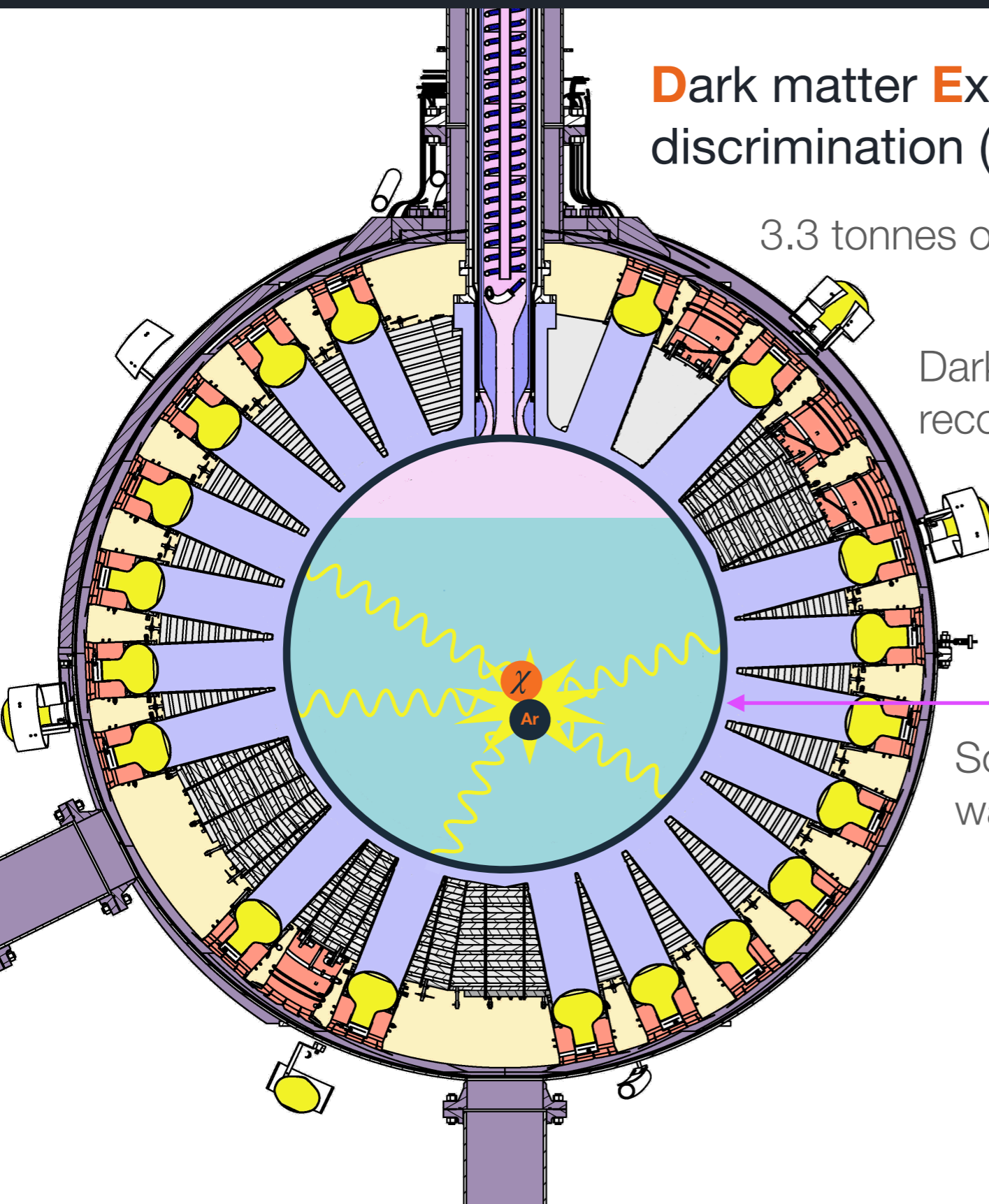
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Scintillation photons pass through TPB wavelength shifter, become 420 nm photons



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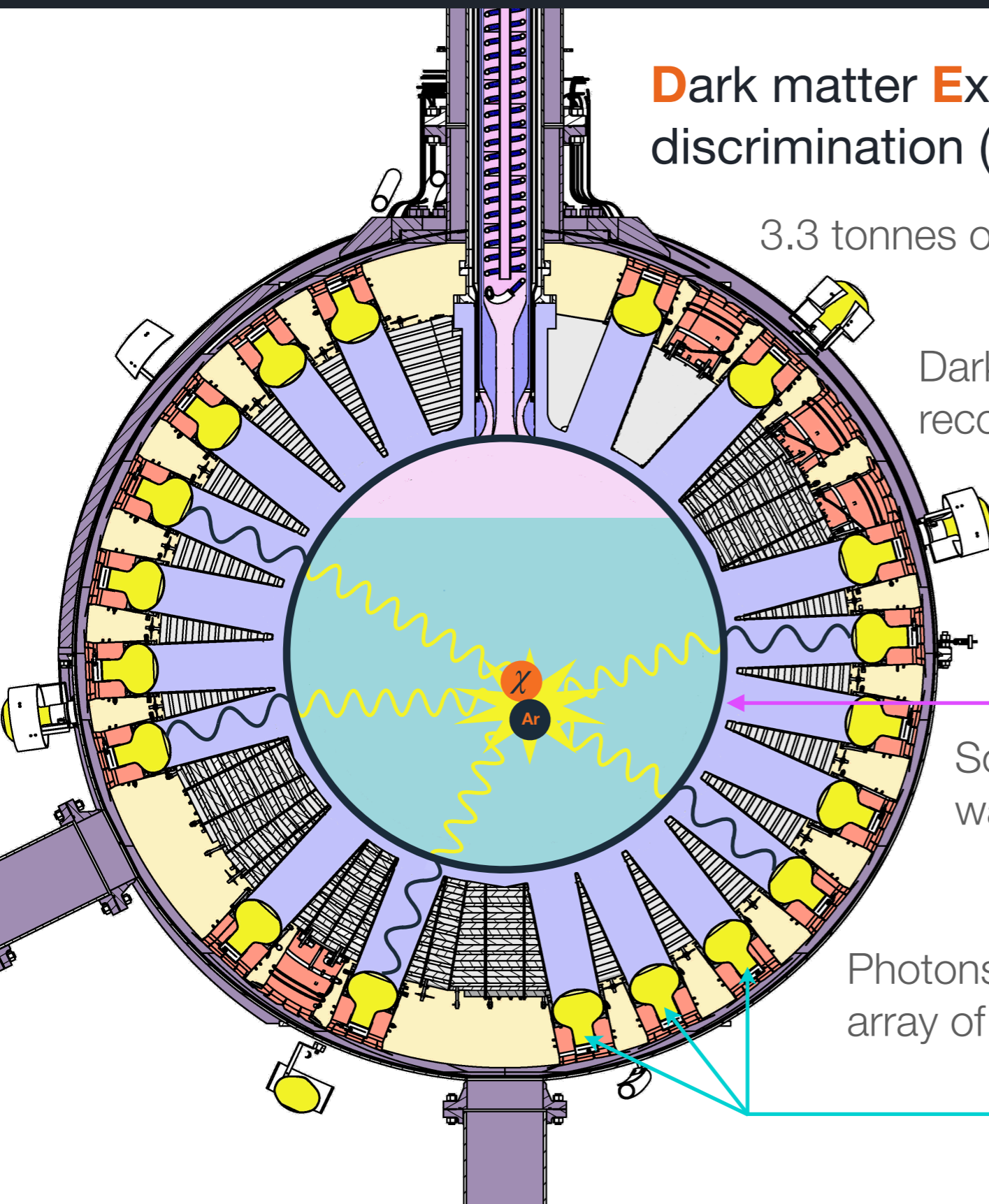
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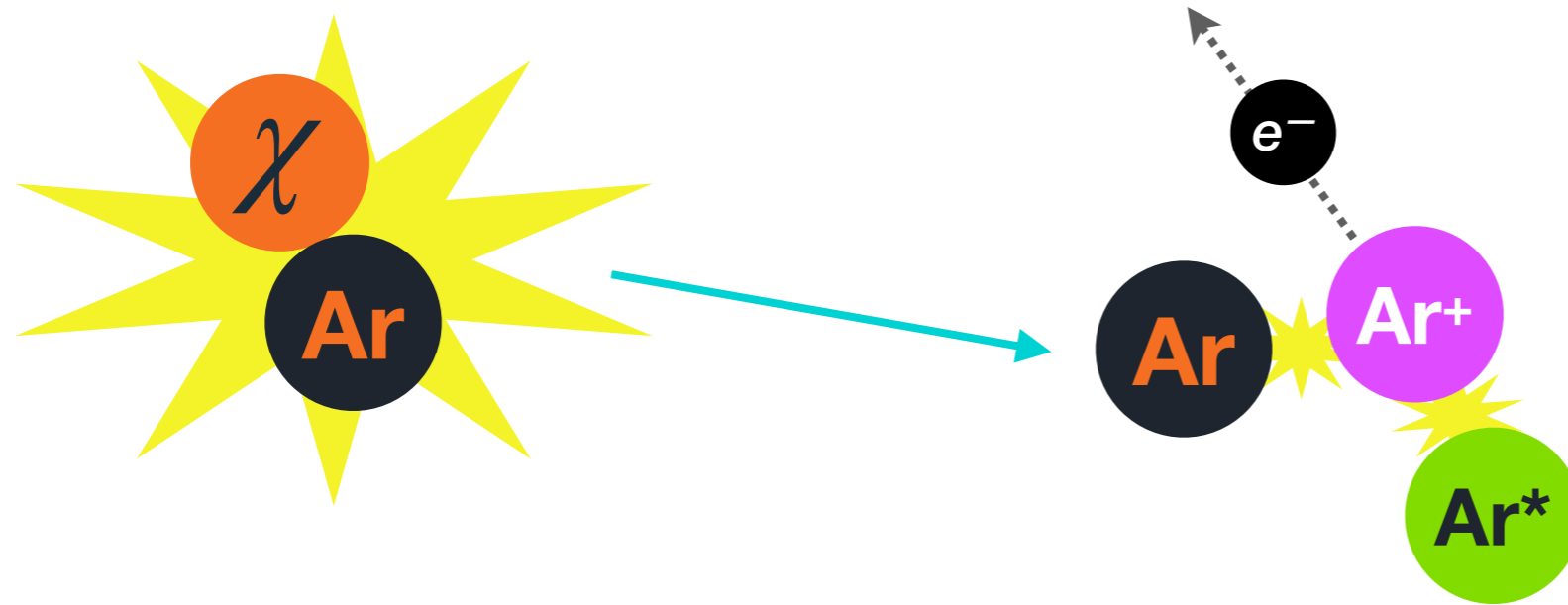
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Photons collected by light guides, detected by array of 255 photomultiplier tubes (PMTs)



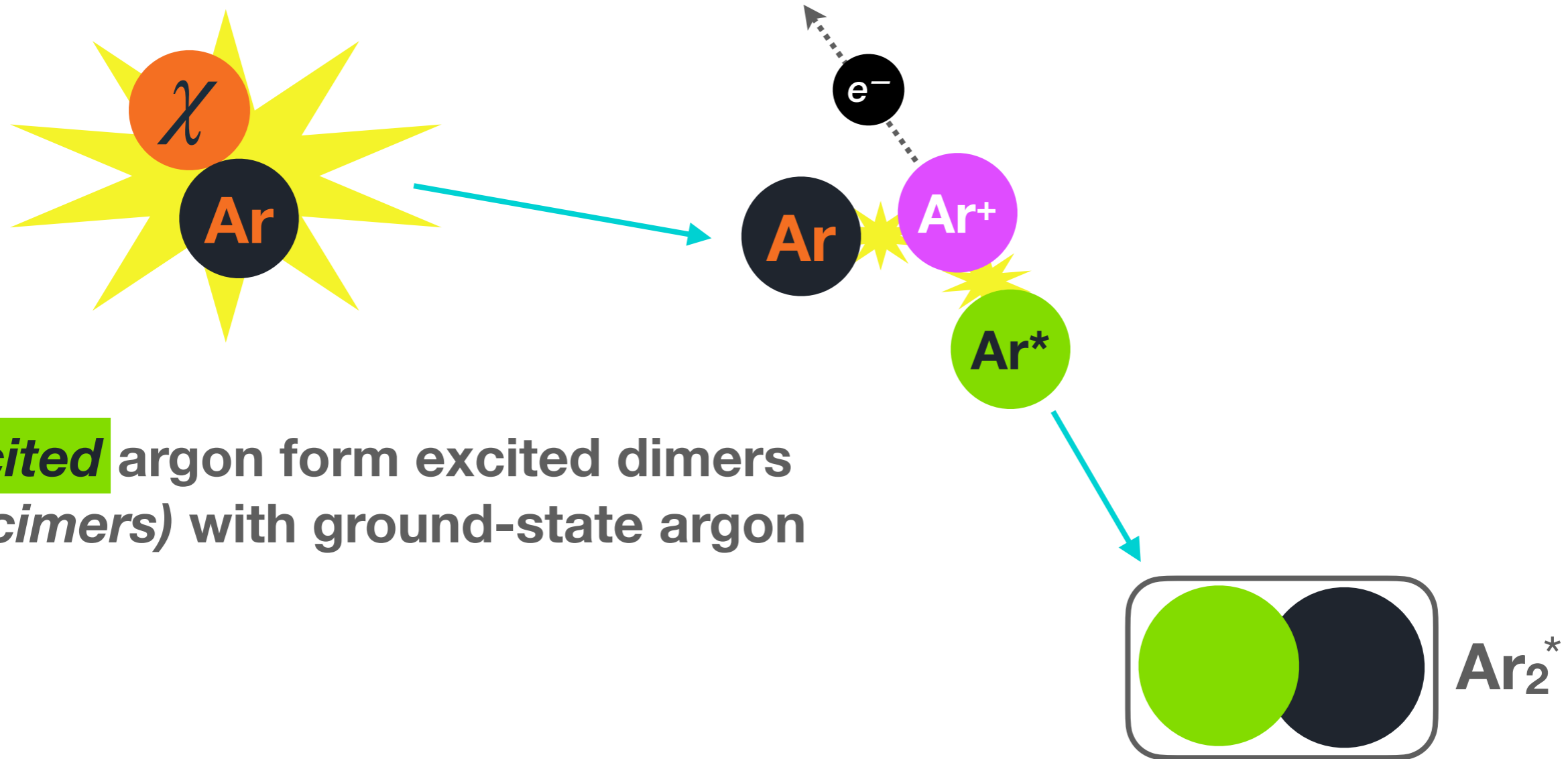


# Overview of DEAP-3600: Liquid Argon Scintillation



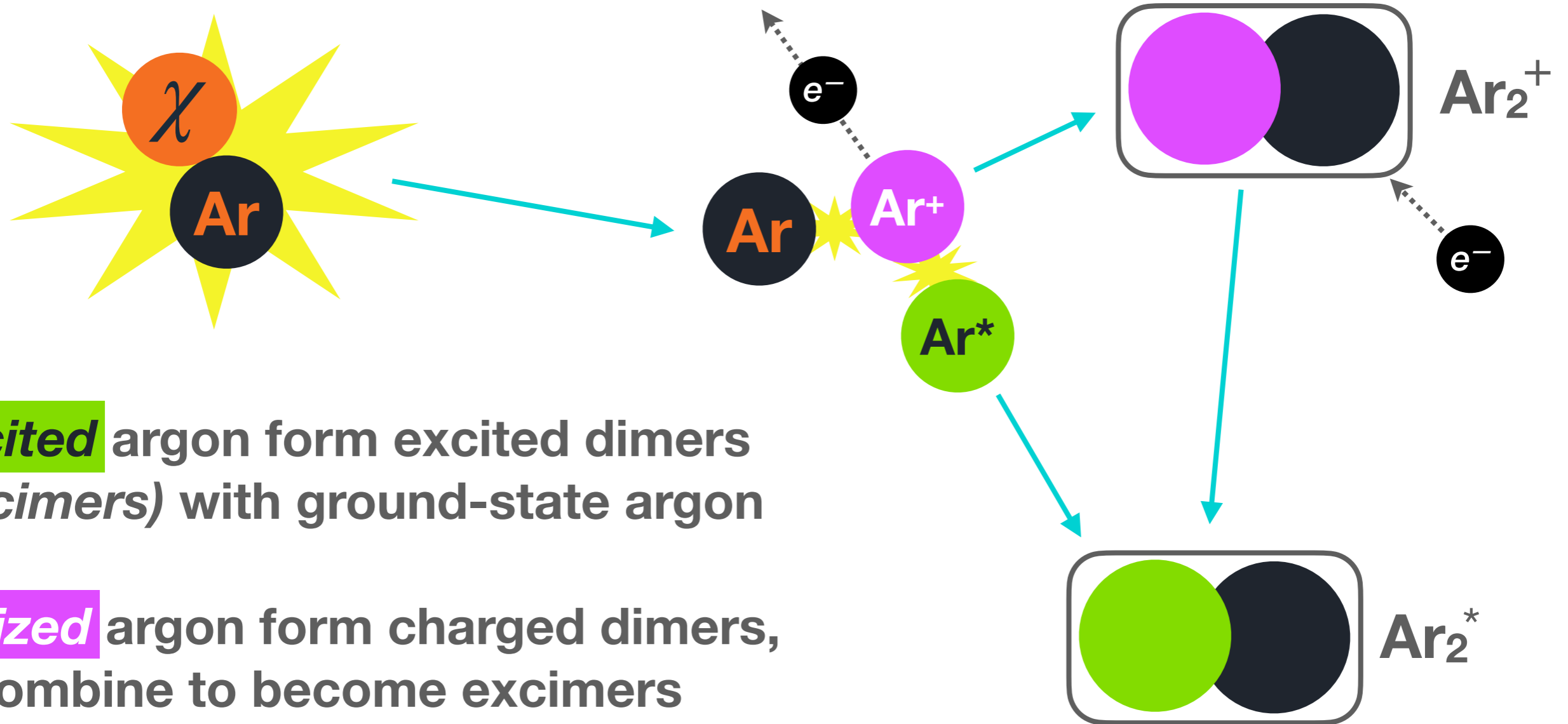
Other atoms in the track of the recoiling argon becomes **excited** or **ionized**

# Overview of DEAP-3600: Liquid Argon Scintillation



**Excited** argon form excited dimers (*excimers*) with ground-state argon

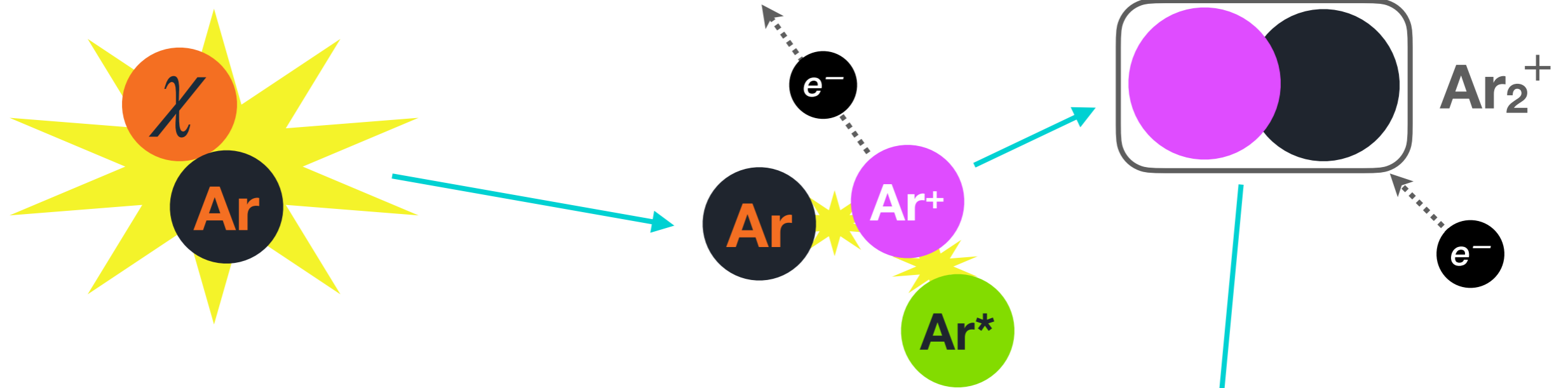
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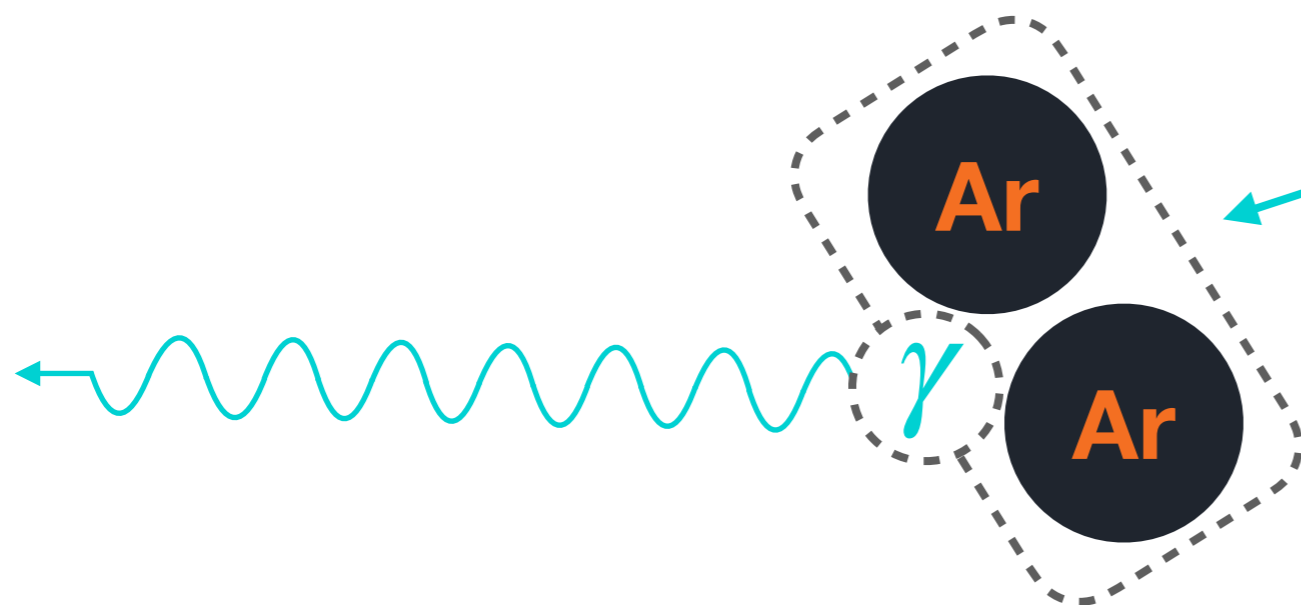
**Excited** argon form excited dimers (*excimers*) with ground-state argon

**Ionized** argon form charged dimers, recombine to become excimers

# Overview of DEAP-3600: Liquid Argon Scintillation



Excimers relax and decay  
producing scintillation photons



**Singlet state**  
 $\tau \approx 7 \text{ ns}$

**Triplet state**  
 $\tau \approx 1.6 \mu\text{s}$

# Overview of DEAP-3600: Pulse Shape Discrimination



## Nuclear Recoils

Scattering directly with argon nuclei; excimers mostly populate the **singlet state**, relax quickly. Induced by:

- Neutrons
- Alphas
- WIMPs

## Electronic Recoils

Scattering with argon atomic electrons, ionizing argon; excimers tend to populate **triplet state**, relax slowly. Induced by:

- Betas (especially  $^{39}\text{Ar}$  at  $\sim 3$  kHz)
- Gammas

# Overview of DEAP-3600: Pulse Shape Discrimination

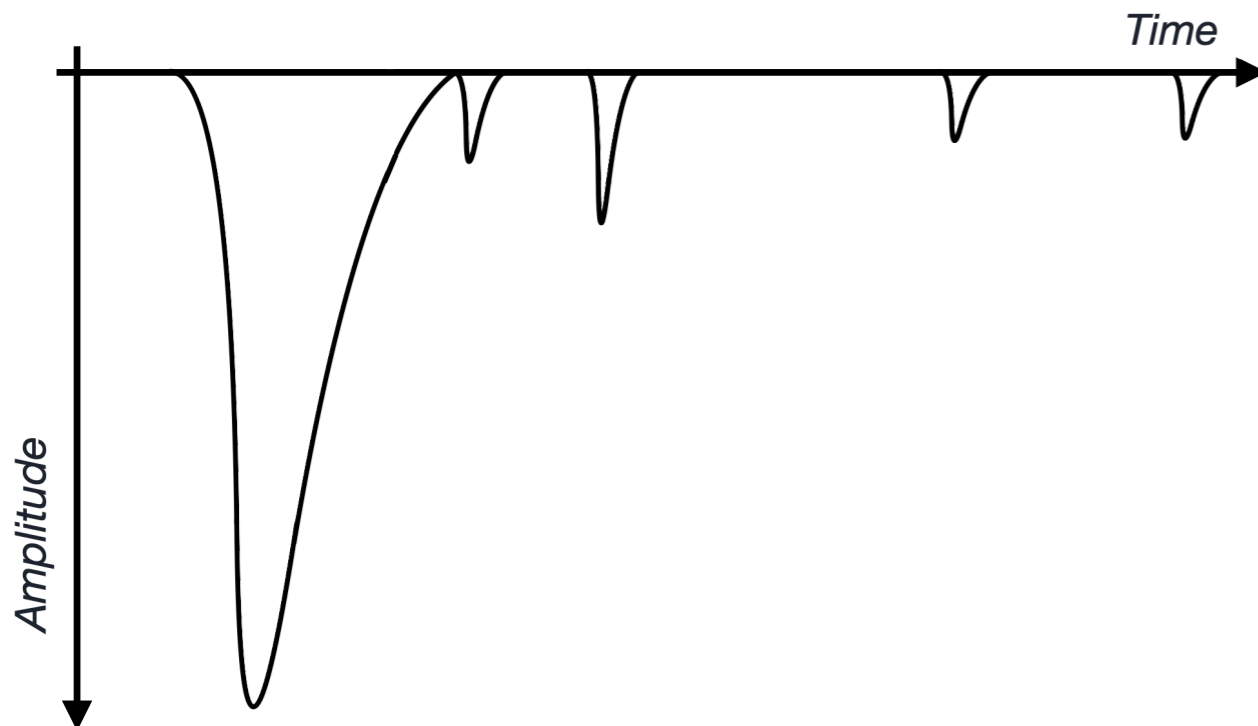


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NR Pulse Shape

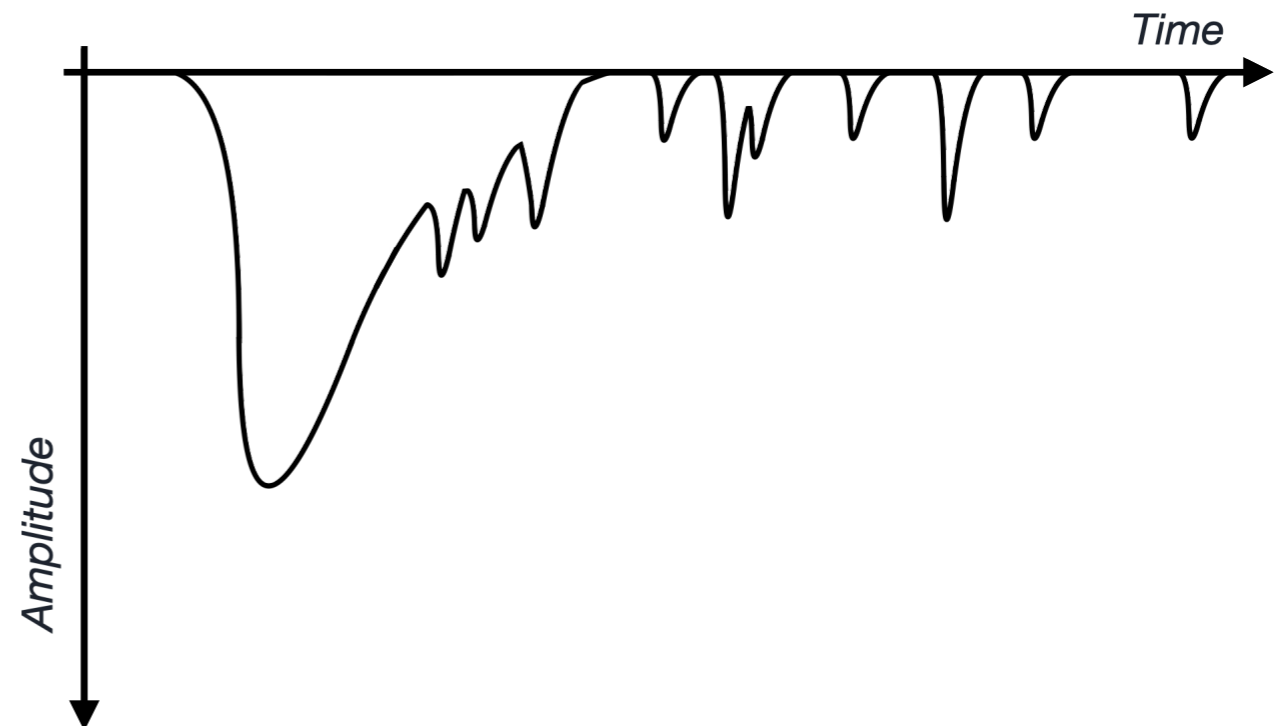


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ER Pulse Shape



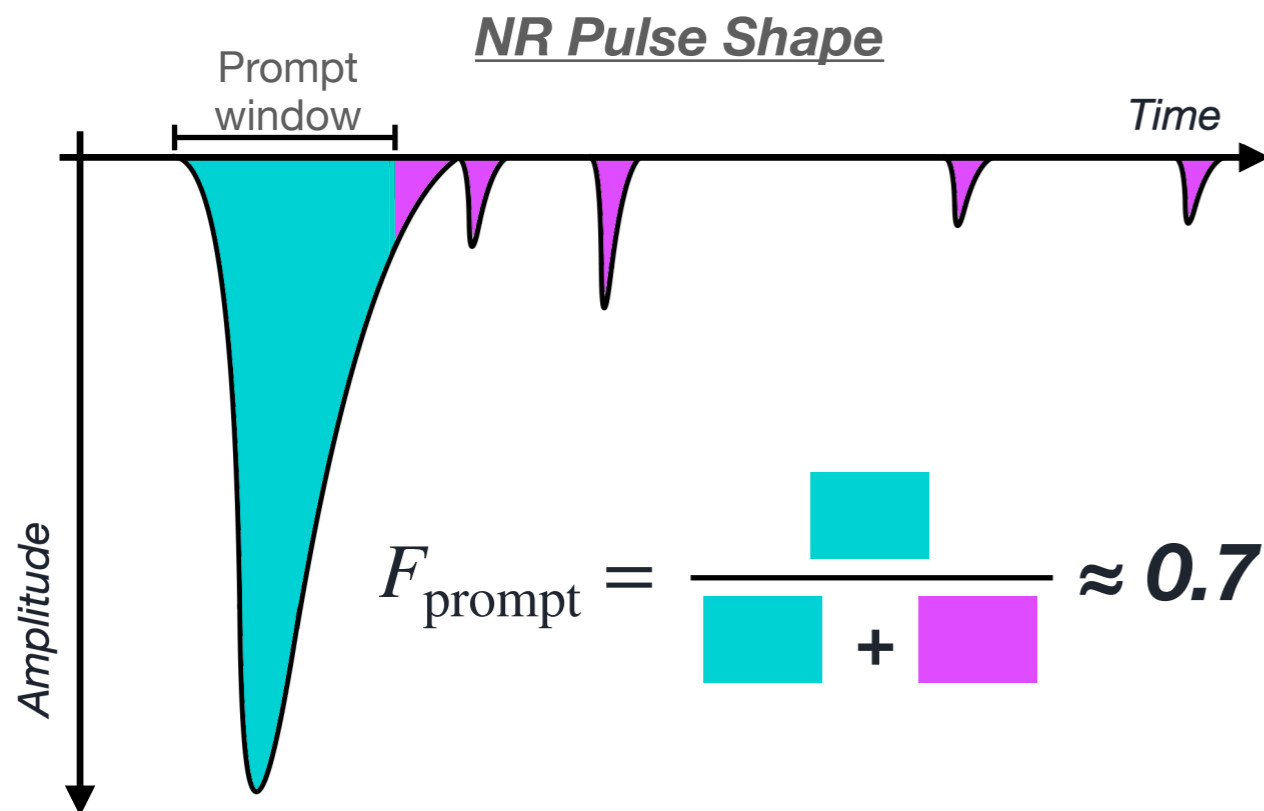
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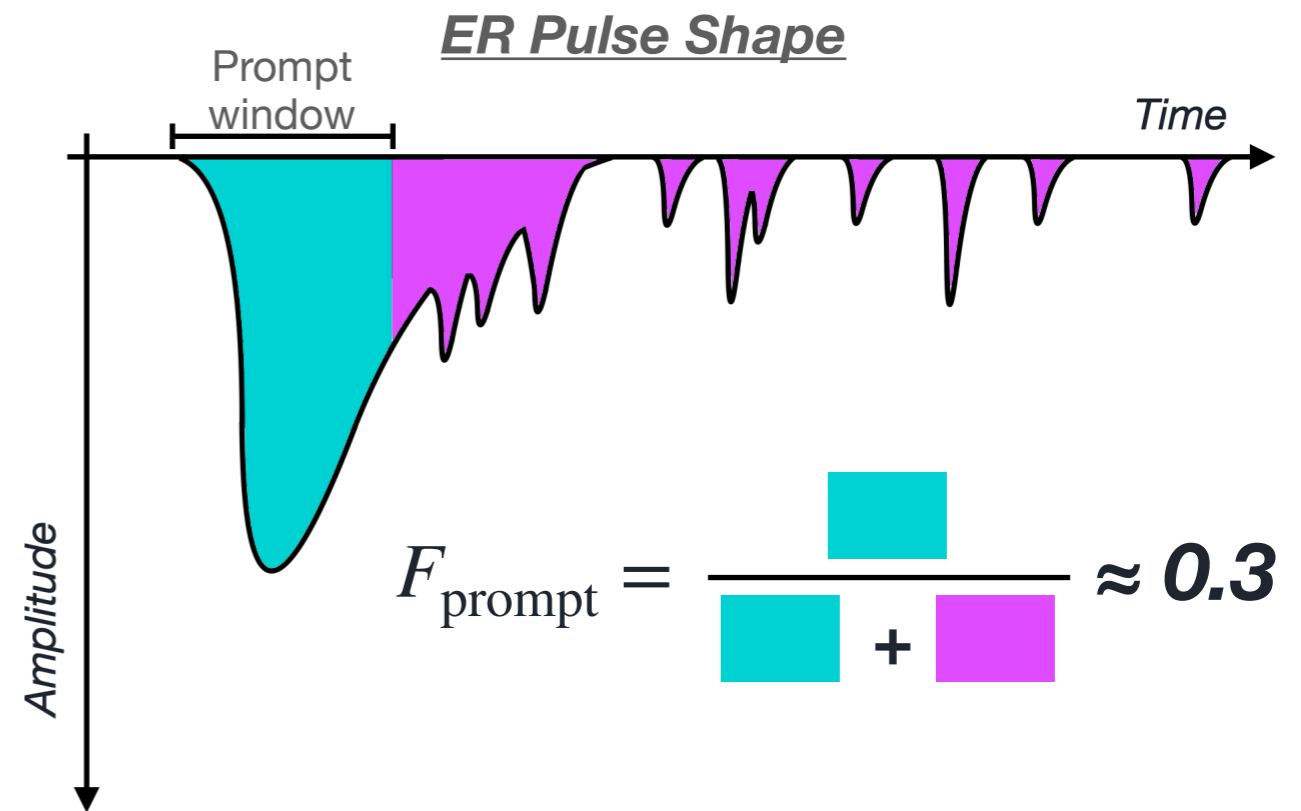
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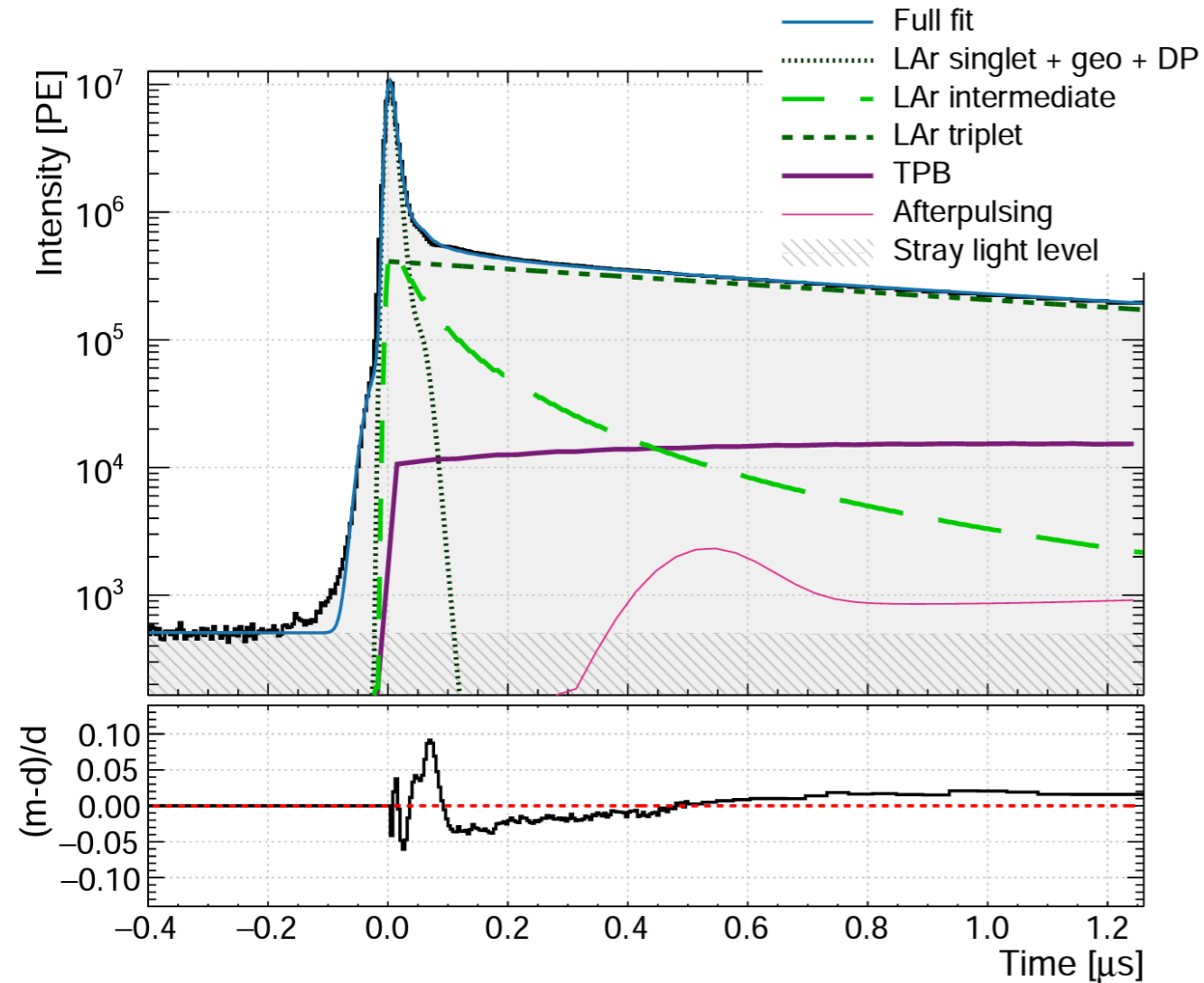
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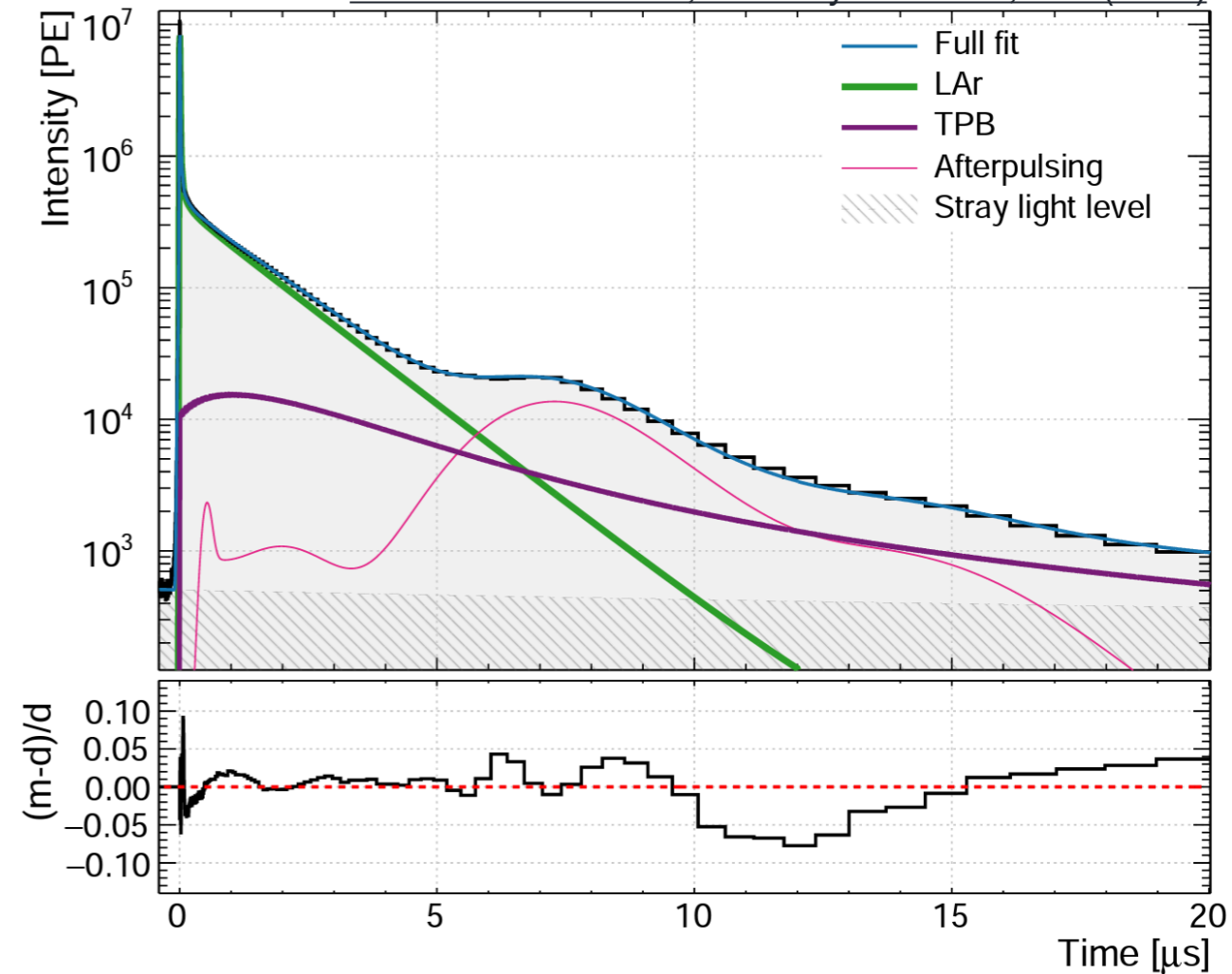
# Precision Measurements: Liquid Argon Pulse Shape



DEAP Collaboration, Eur. Phys. J. C 80, 303 (2020)



DEAP Collaboration, Eur. Phys. J. C 80, 303 (2020)



- Characterized the LAr pulse shape, accounting for detector geometry, and contributions from TPB, PMT afterpulsing, double/late pulsing, and stray light
- Pulse shape includes known singlet & triplet components, also *intermediate* component

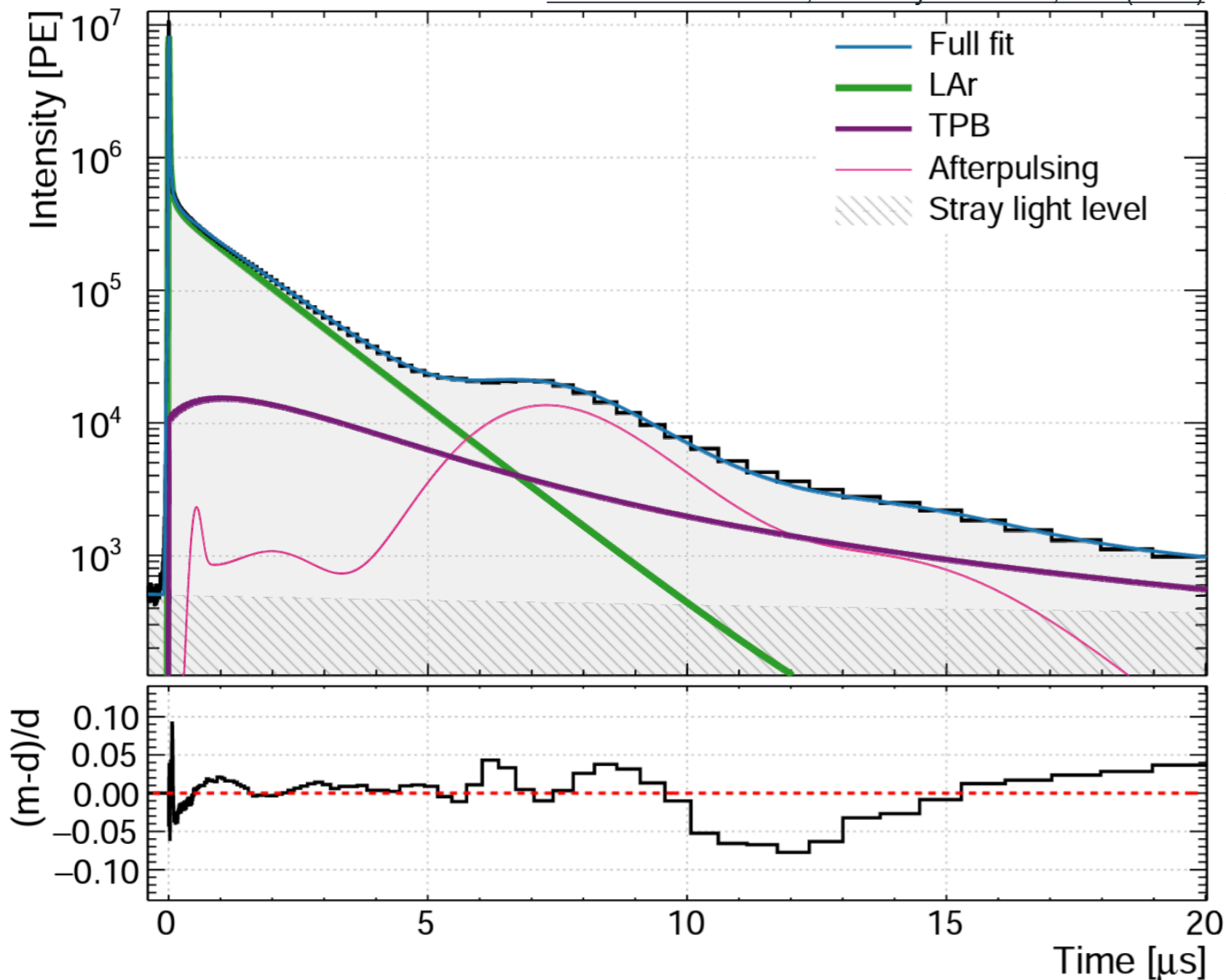


# Precision Measurements: PSD in 4.5 Tonne-Year Exposure



- Detailed pulse shape modelling allows for ability to separate scintillation from PMT artifacts; i.e. PMT afterpulsing

DEAP Collaboration, Eur. Phys. J. C 80, 303 (2020)



**Prompt Window Length**

$$F_{\text{prompt}} = \frac{\sum_{t=-28 \text{ ns}}^{60 \text{ ns}} \text{PE}(t)}{\sum_{t=-28 \text{ ns}}^{10 \mu\text{s}} \text{PE}(t)}$$

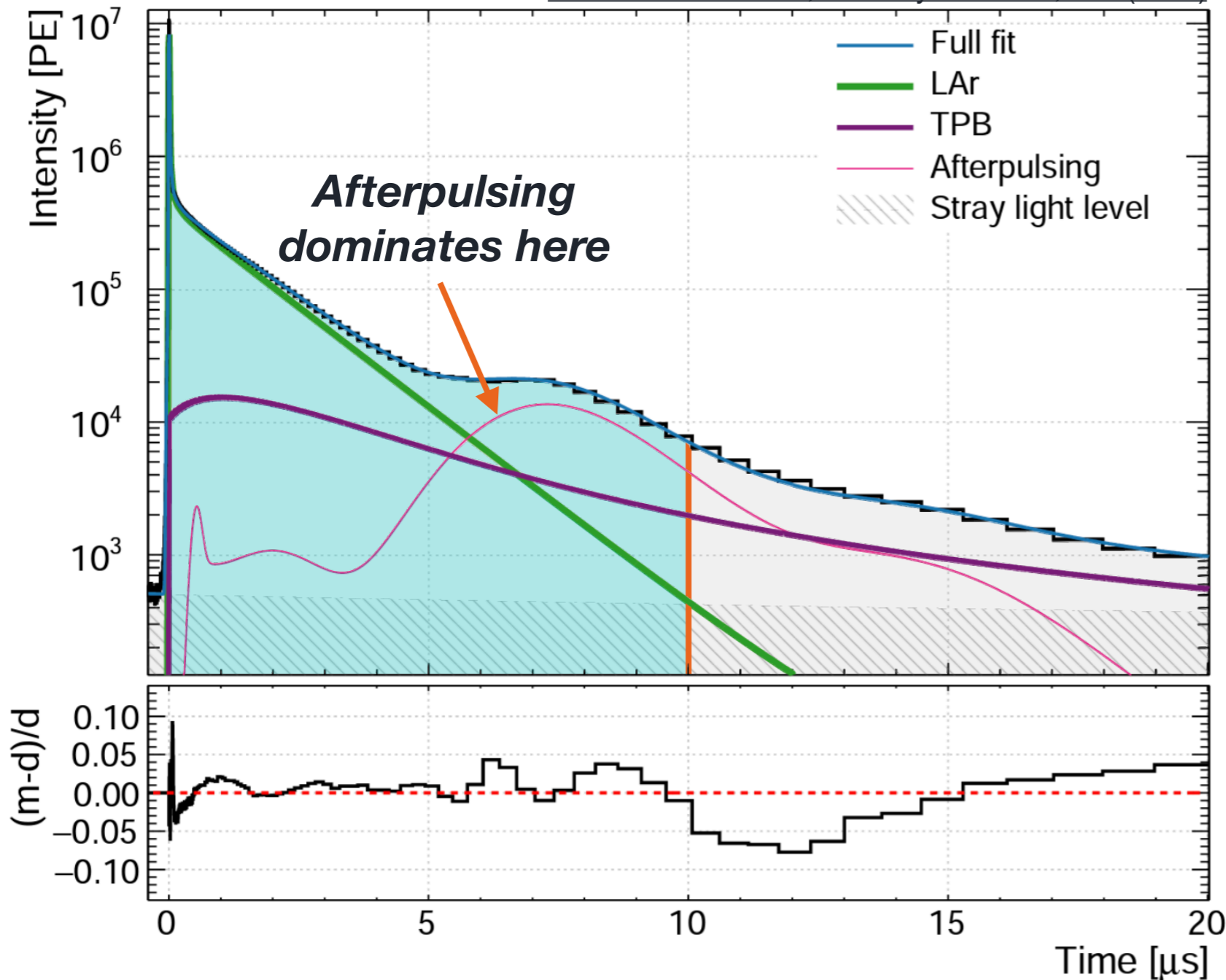
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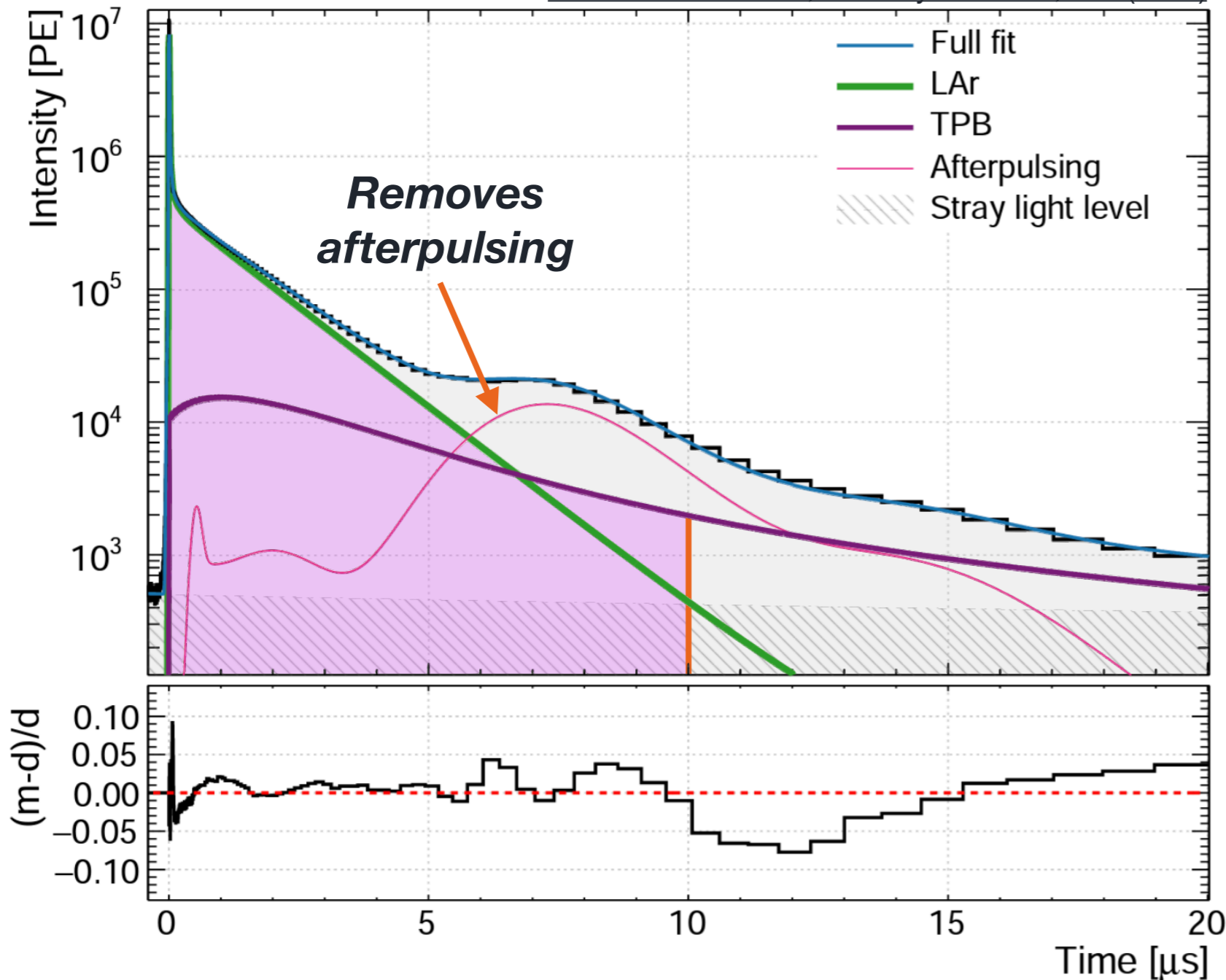
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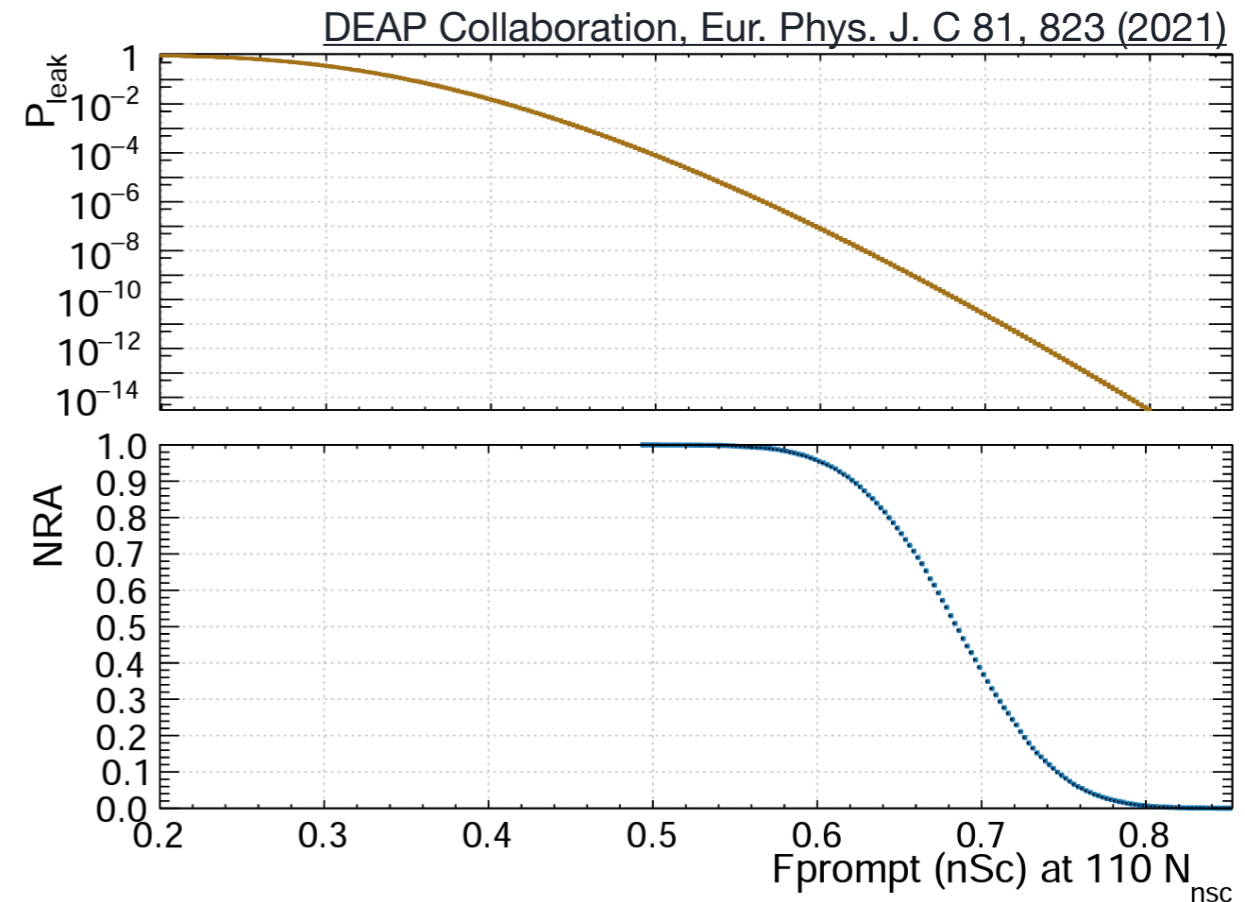
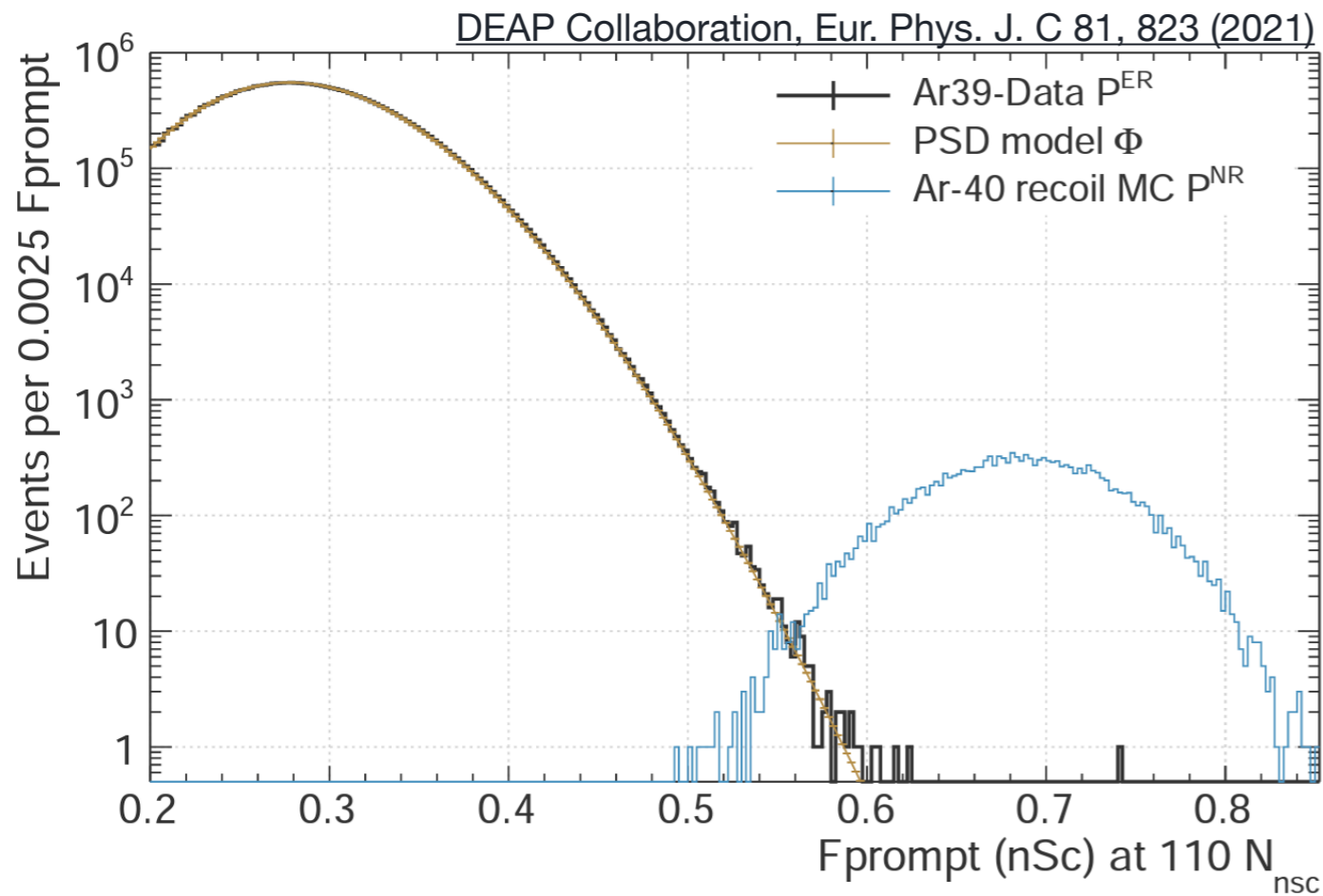
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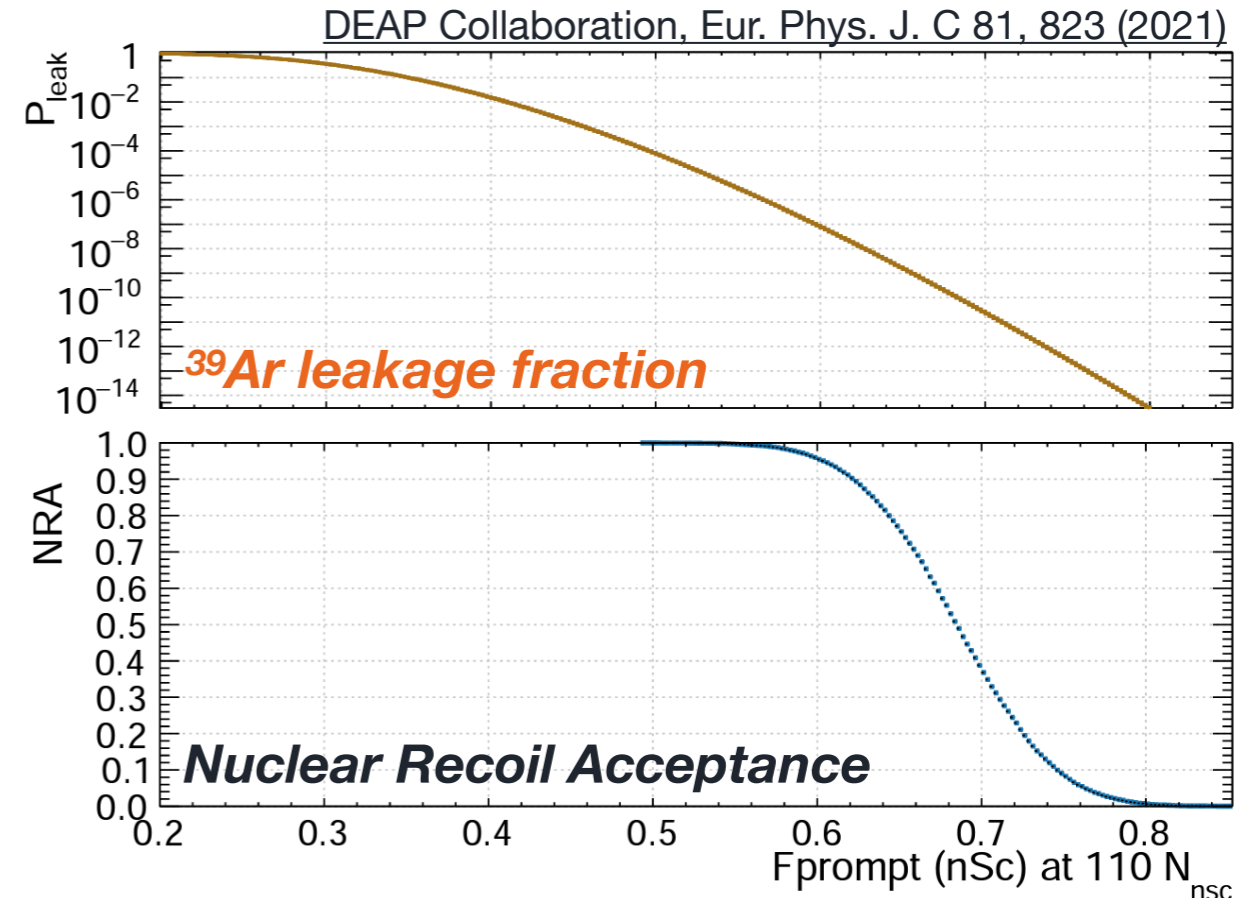
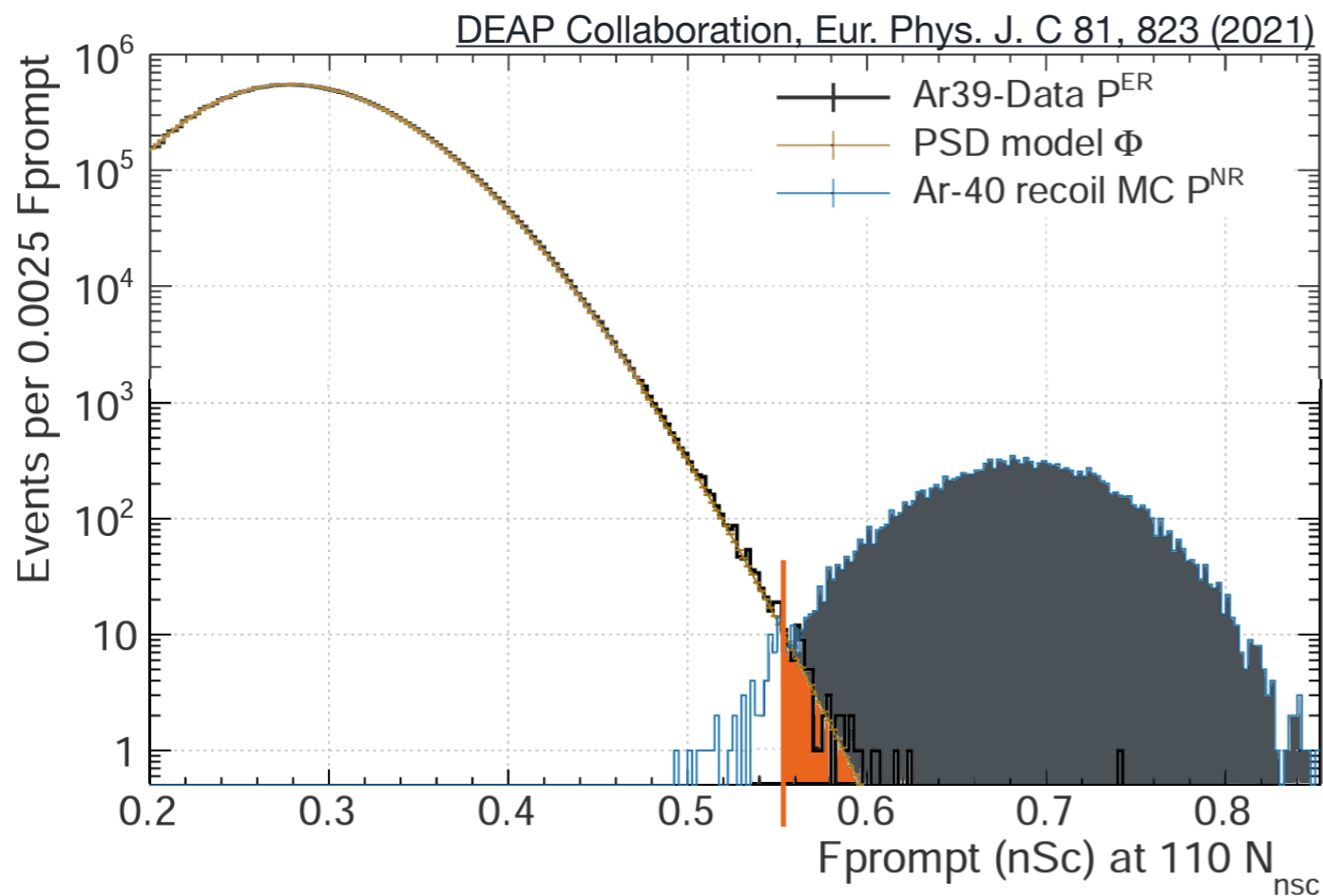
$$\text{PE} \equiv \text{Bayesian Photoelectron Counting}$$

# Precision Measurements: PSD in 4.5 Tonne-Year Exposure



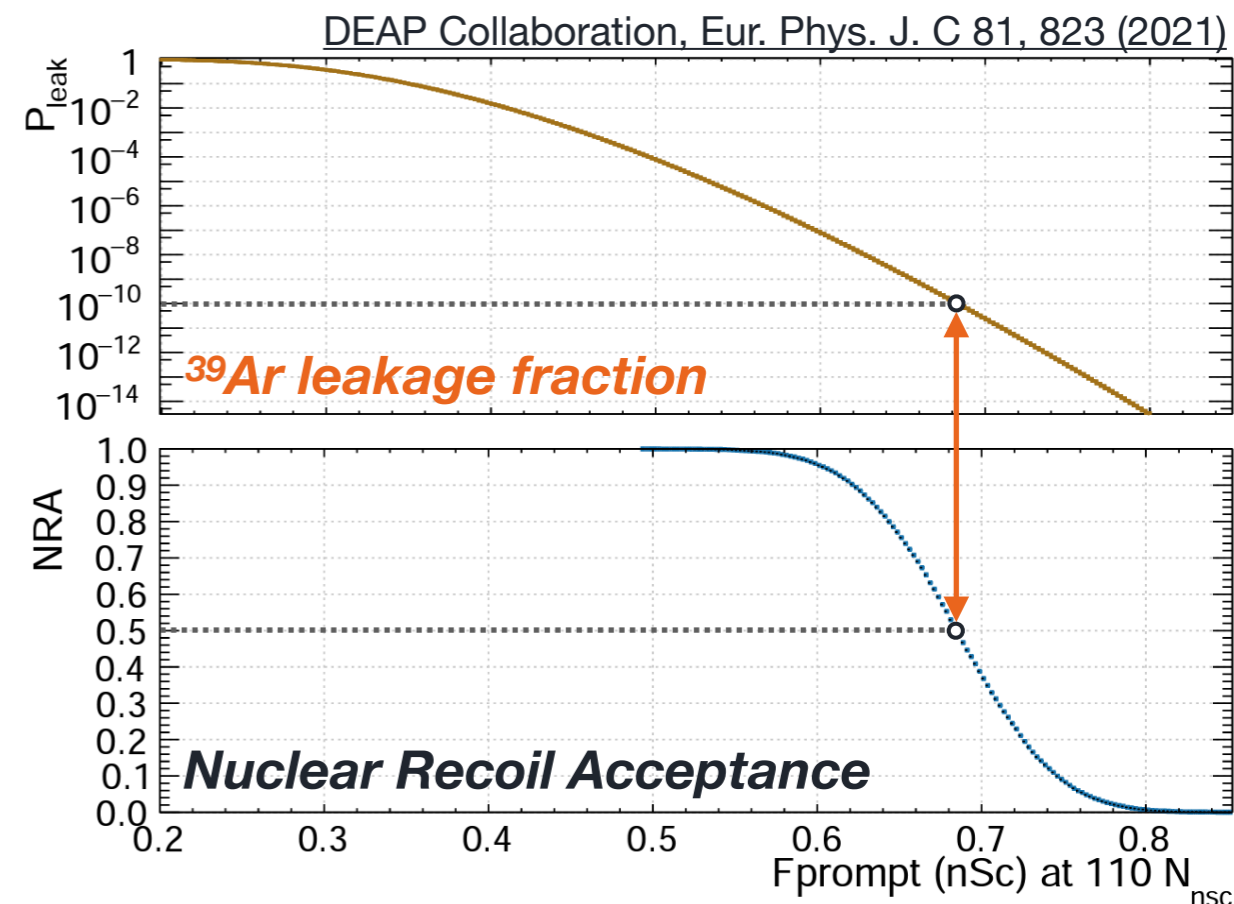
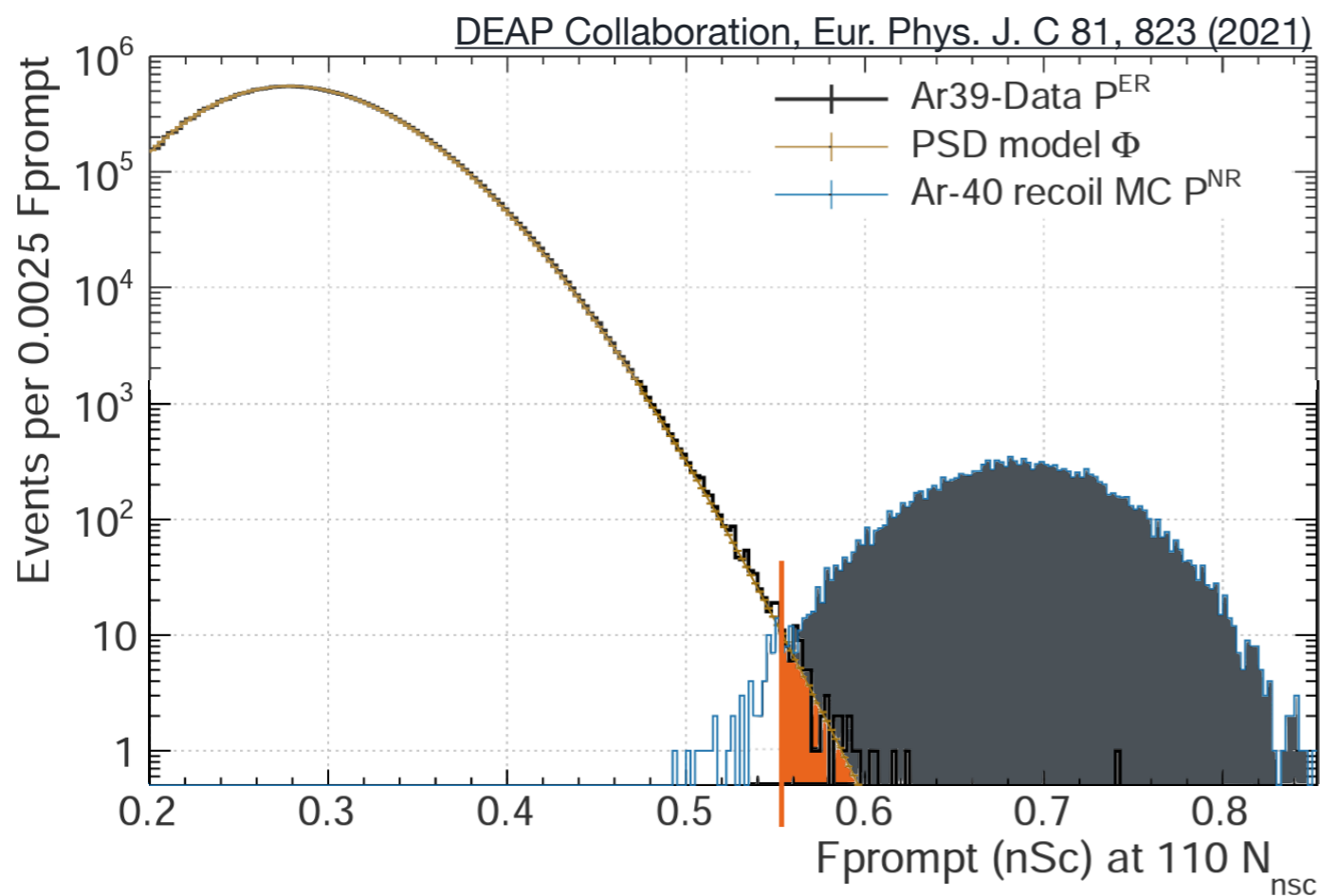
- PSD model tested with both energy estimators: total integrated charge & with afterpulsing removal

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- $^{39}\text{Ar}$  leakage is reduced by an order magnitude with afterpulsing removal compared to total charge integration

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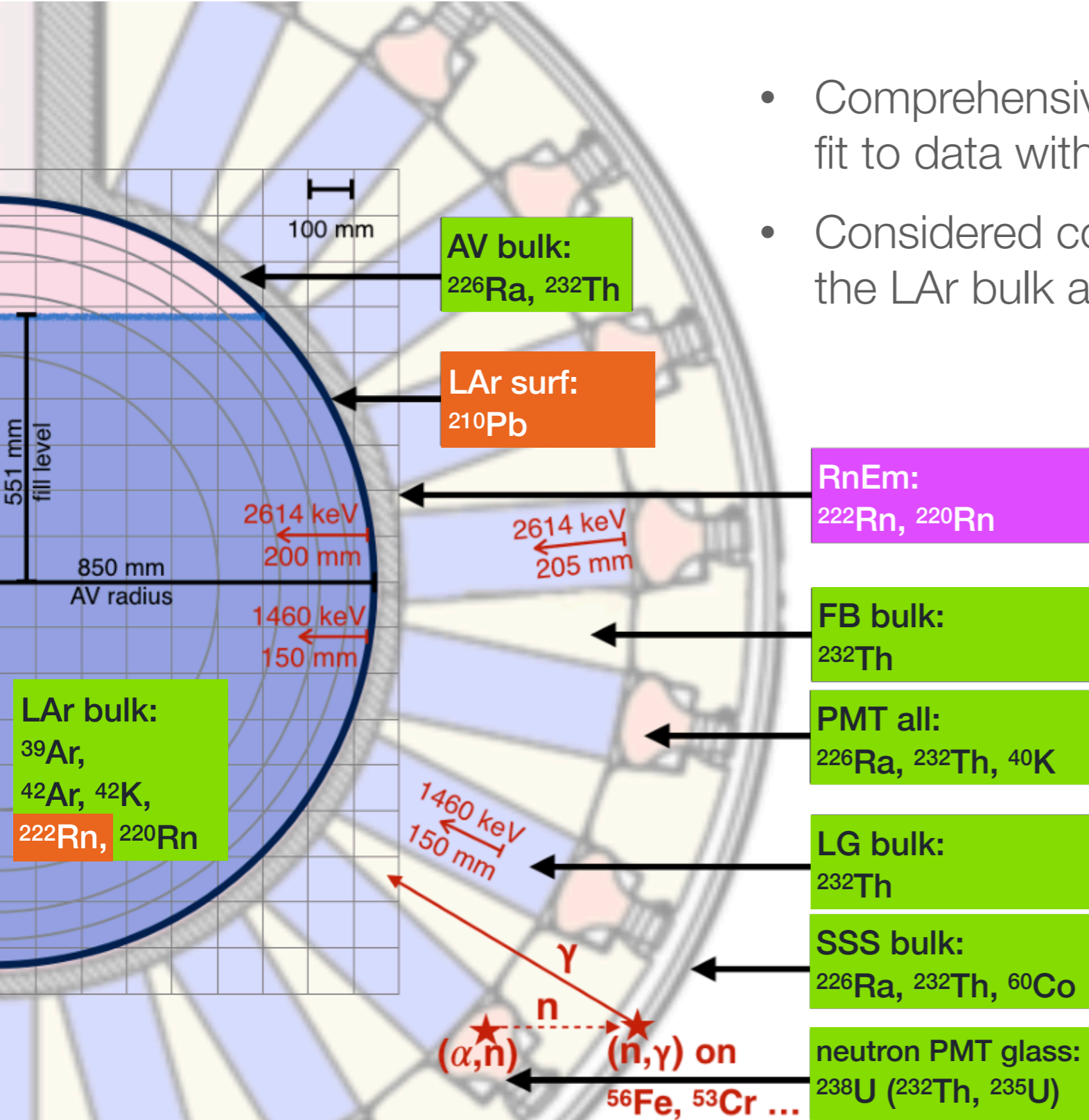


- PSD model tested with both energy estimators: total integrated charge & with afterpulsing removal
- $^{39}\text{Ar}$  leakage is reduced by an order magnitude with afterpulsing removal compared to total charge integration
- Result: **world leading PSD!**  
 $10^{-10}$  leakage fraction of  $^{39}\text{Ar}$  for 50% NR acceptance at 110 PE (117.5 keV<sub>ee</sub>)

# Precision Measurements: Electromagnetic Backgrounds



DEAP Collaboration, Phys. Rev. D 100, 072009



- Comprehensive electromagnetic backgrounds model fit to data with BAT (Bayesian Analysis Toolkit)
- Considered components include sources located in the LAr bulk all the way out to the stainless steel shell

## EM Backgrounds Model

*14 sources as free parameters*

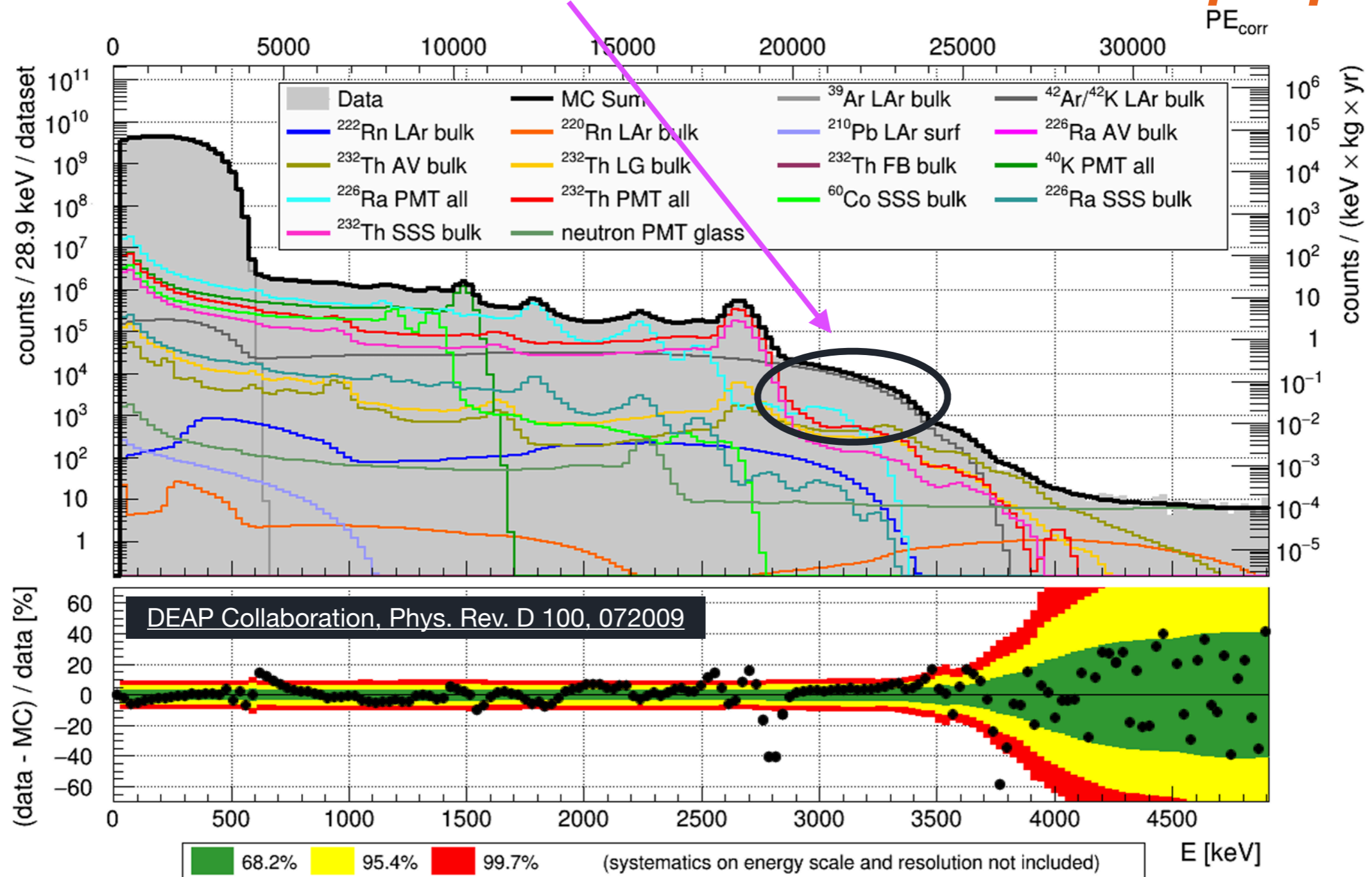
*2 sources as fixed parameters*

*2 sources included as systematics*

# Precision Measurements: Electromagnetic Backgrounds

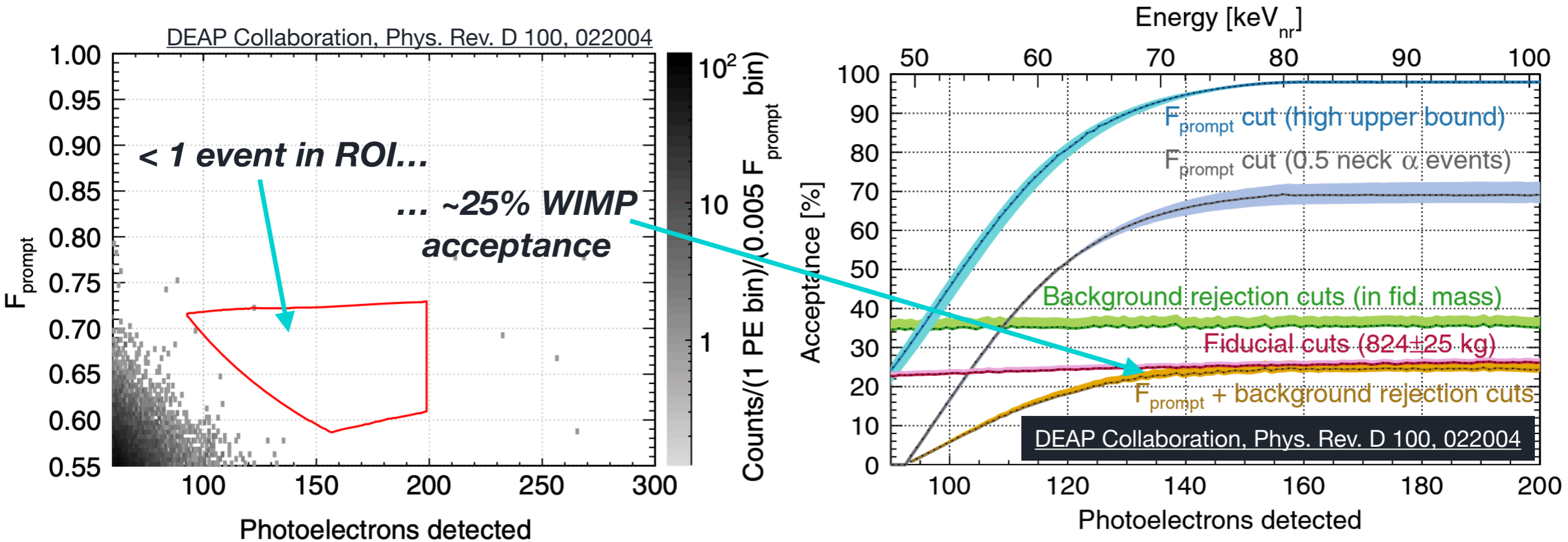


- $^{42}\text{Ar}$  betas are source of background for DarkSide-20k, GERDA, LEGEND—previously measurements of its specific activity are in tension, have large uncertainties
- DEAP measures  $^{42}\text{Ar}$  activity via  $^{42}\text{K}$  beta decay:  **$A = 40.4 \pm 5.9 \mu\text{Bq/kg}$**



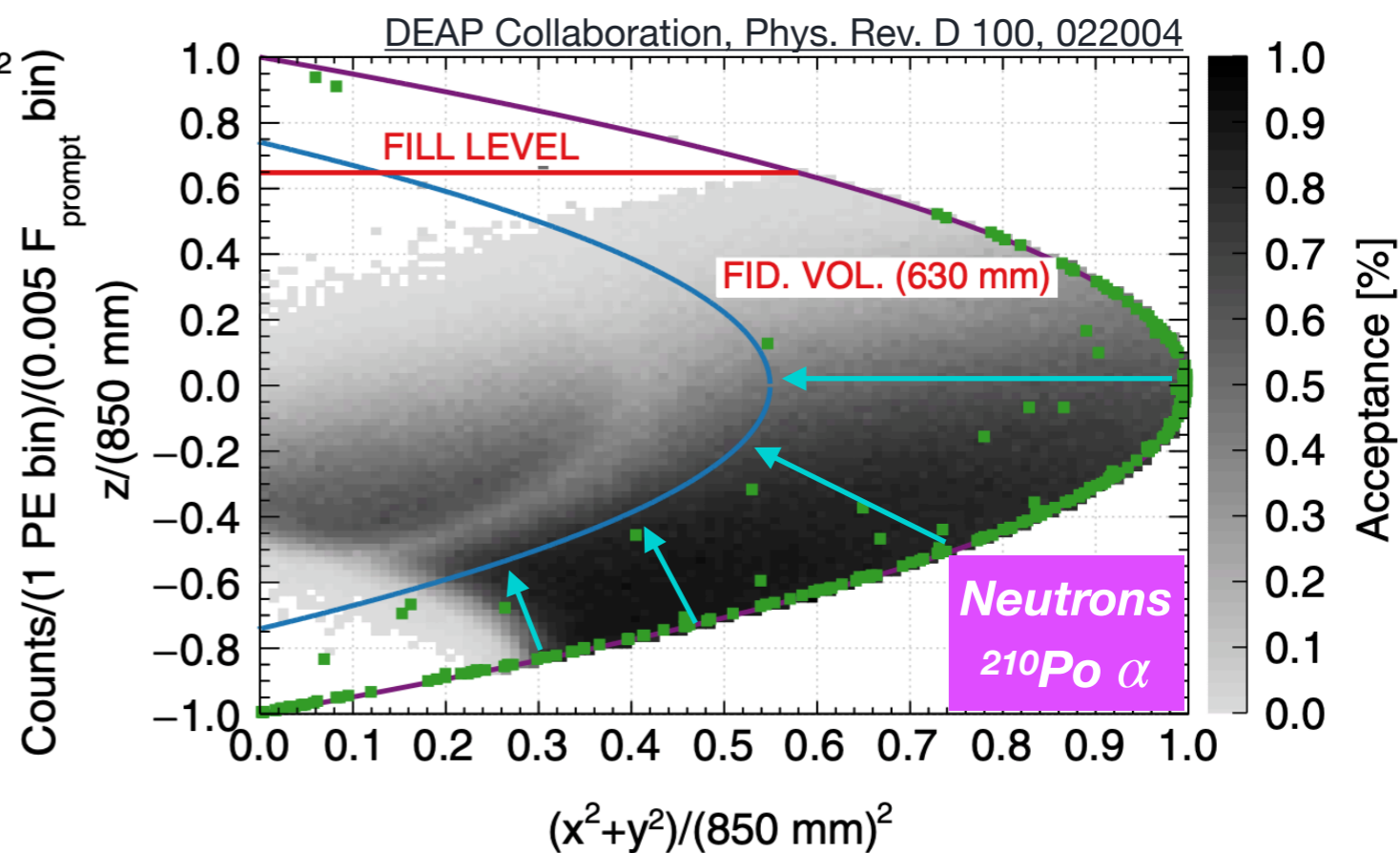
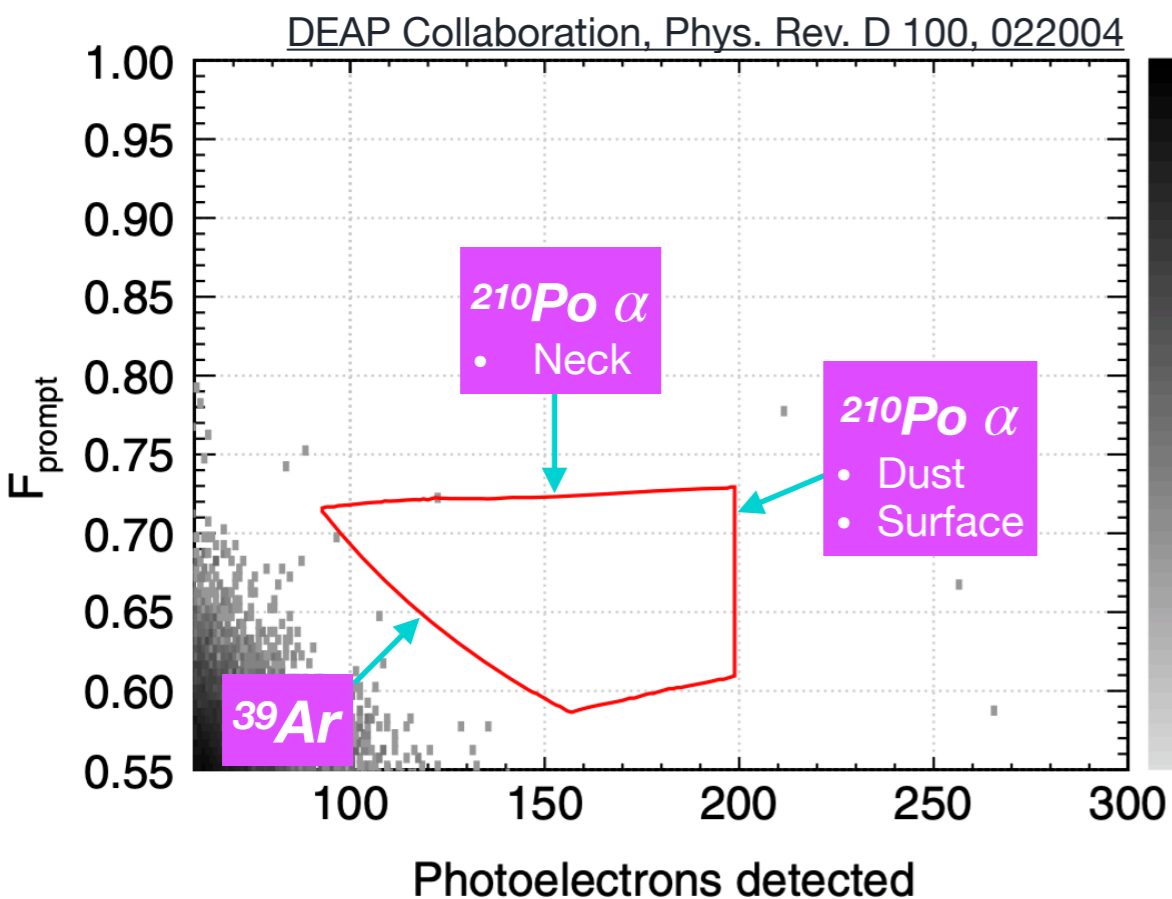


# WIMP Searches: DEAP Standard Analysis



- DEAP's 231 live-day exposure with region of interest (ROI), fiducial volume (FV), and event selection cuts had zero background events

# WIMP Searches: Profile Likelihood Ratio



- DEAP's 231 live-day exposure with region of interest (ROI), fiducial volume (FV), and event selection cuts had zero background events
- Improved background model and machine learning algorithms will allow us to expand ROI and FV, as well as ease event selection cuts

# WIMP Searches: Profile Likelihood Ratio



## Define a Likelihood Function

$$\mathcal{L}(\mathbf{x} | \sigma, \theta) = \mathcal{L}_{\text{PDFs}}(\mathbf{x} | \sigma, \theta) \cdot \mathcal{L}_{\text{Con}}(\theta) \cdot \mathcal{L}_{\text{SB}}(\theta)$$

Set of observed data points

WIMP-nucleon elastic  
scattering cross-section

Set of nuisance  
parameters (systematics)

# WIMP Searches: Profile Likelihood Ratio



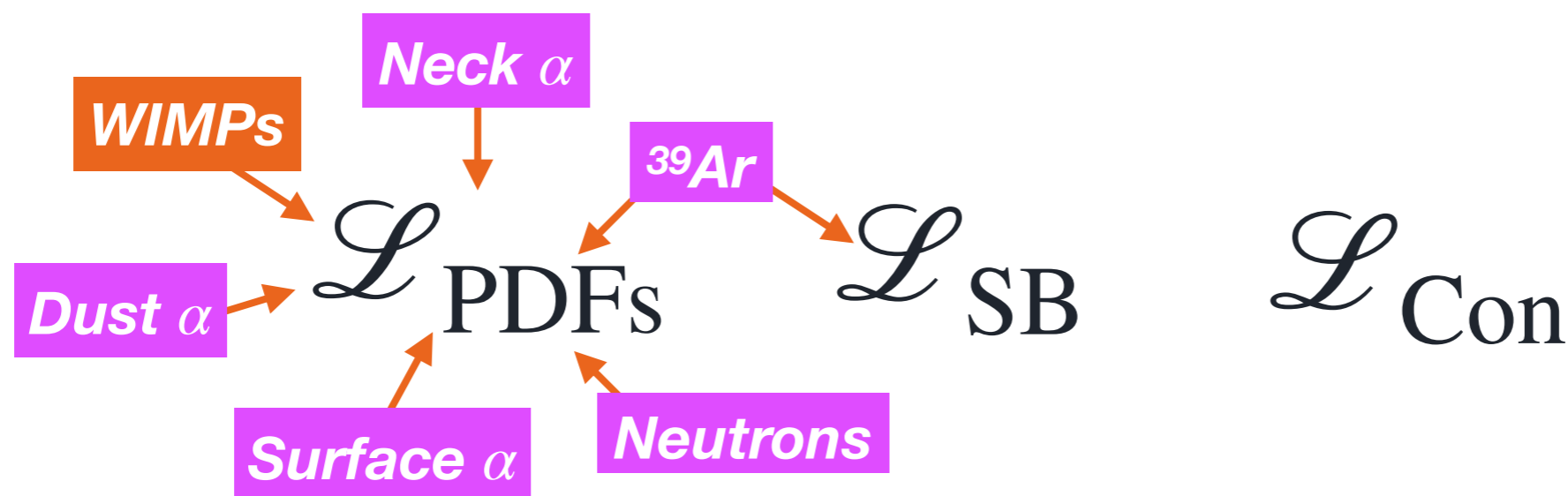
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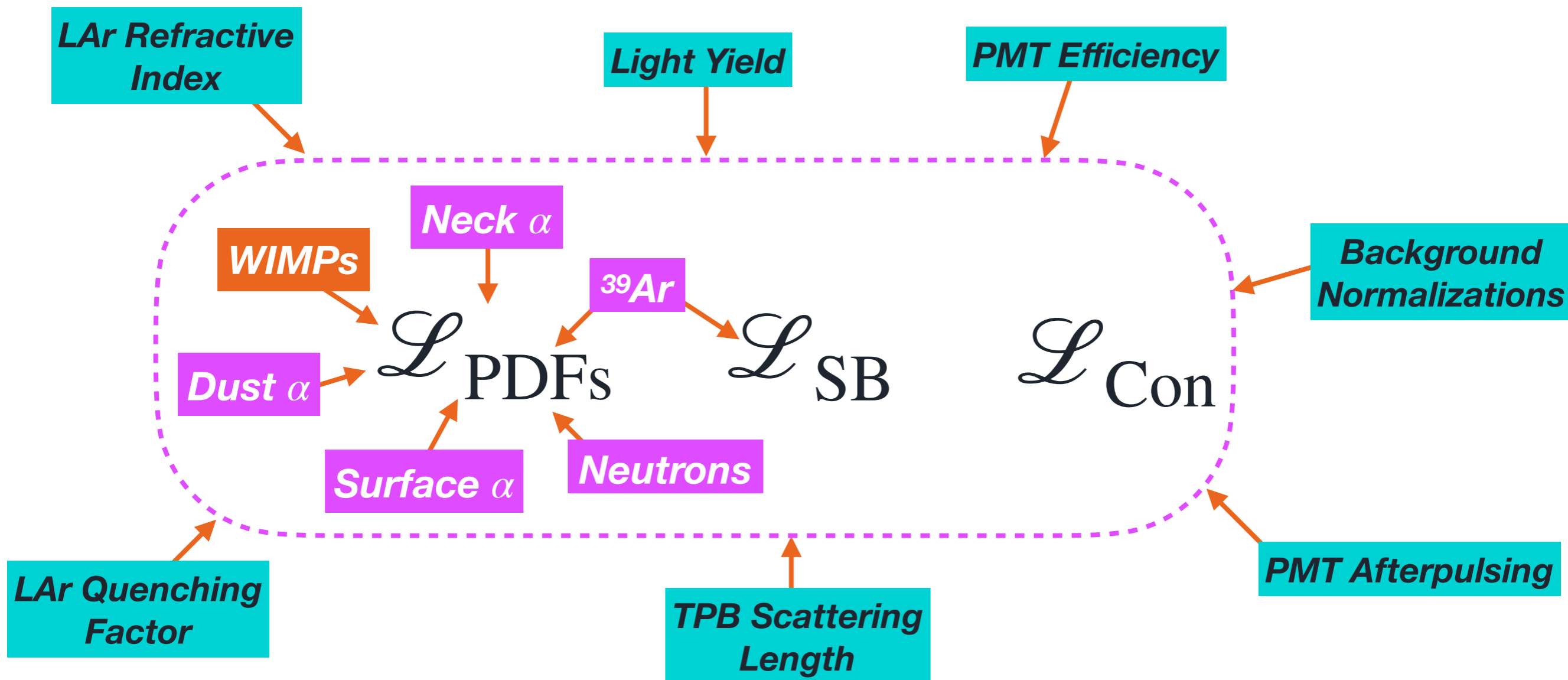
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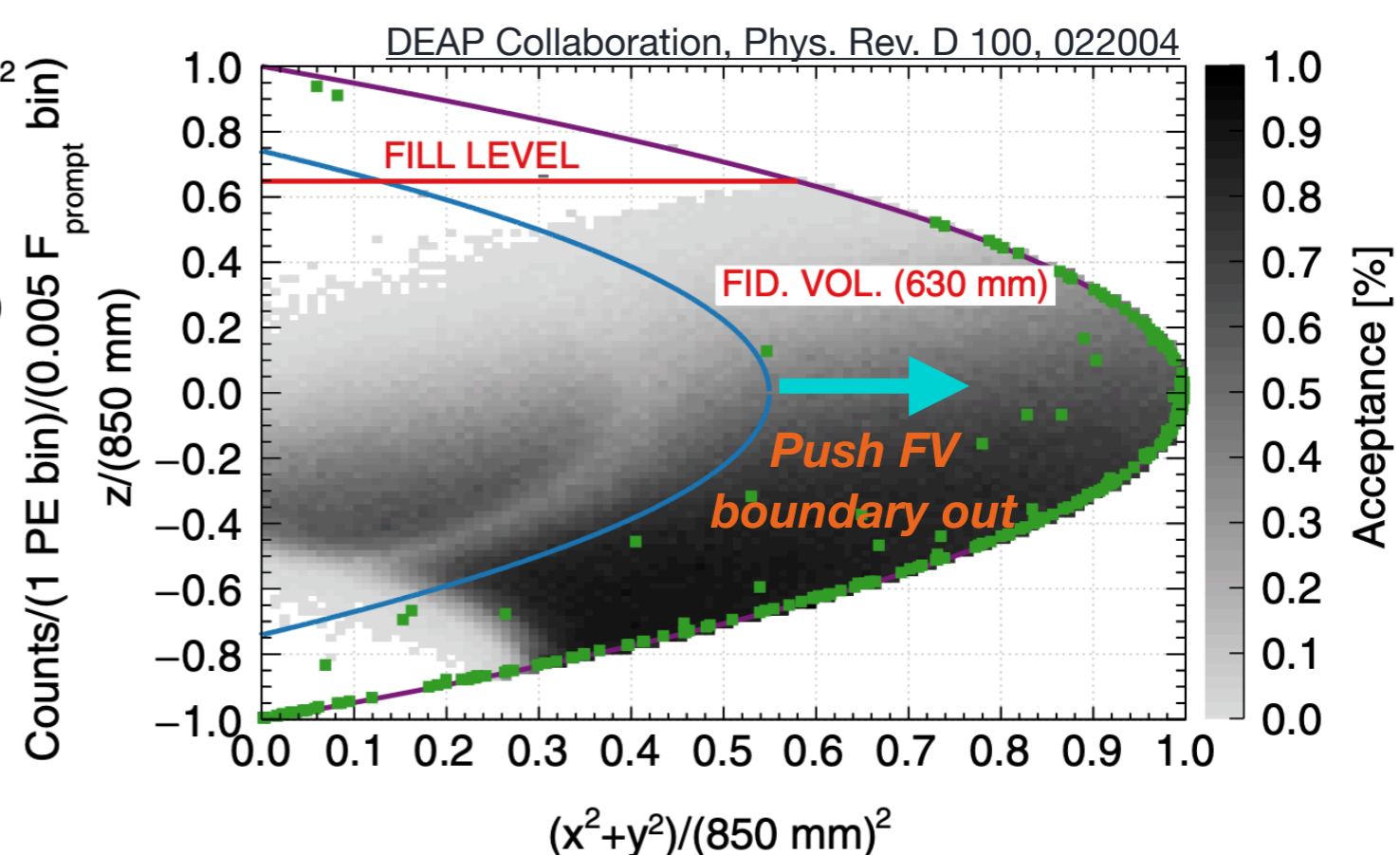
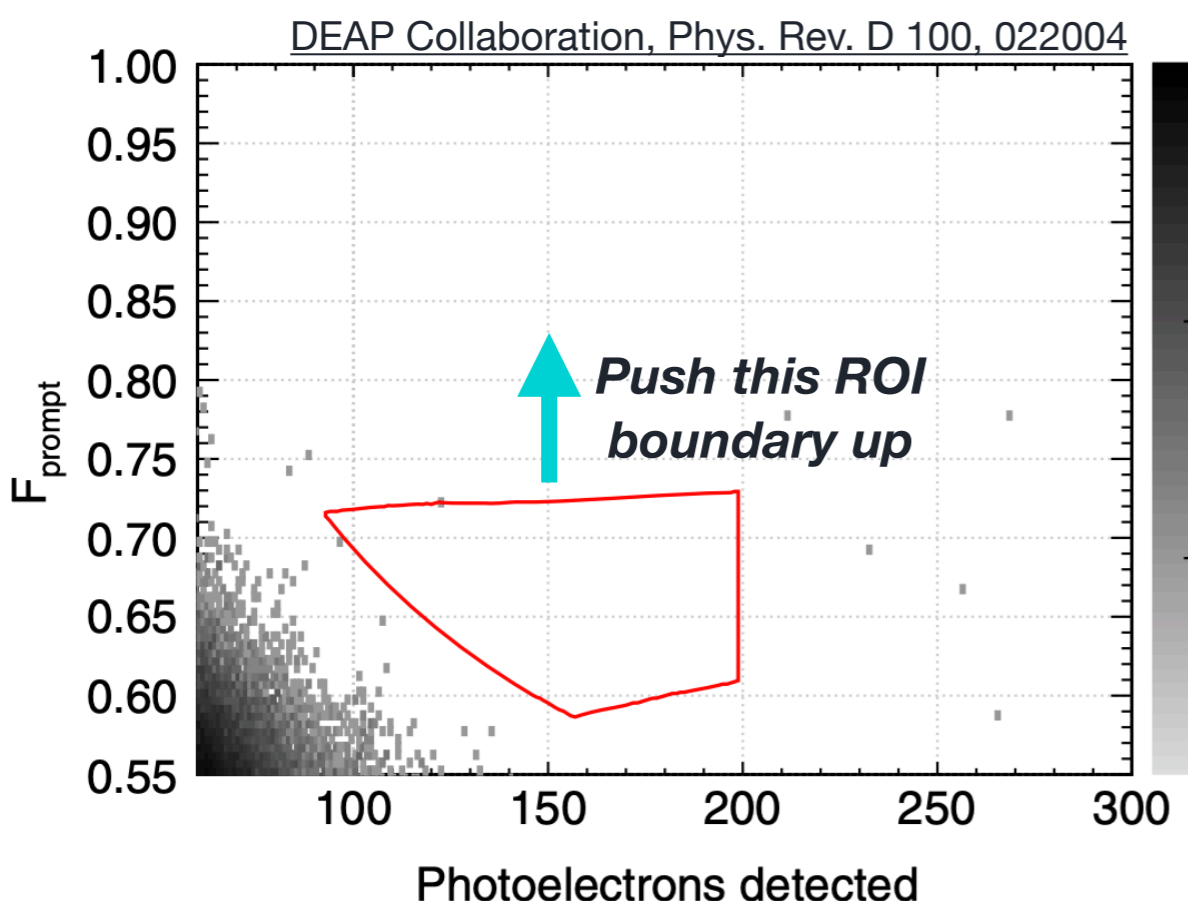
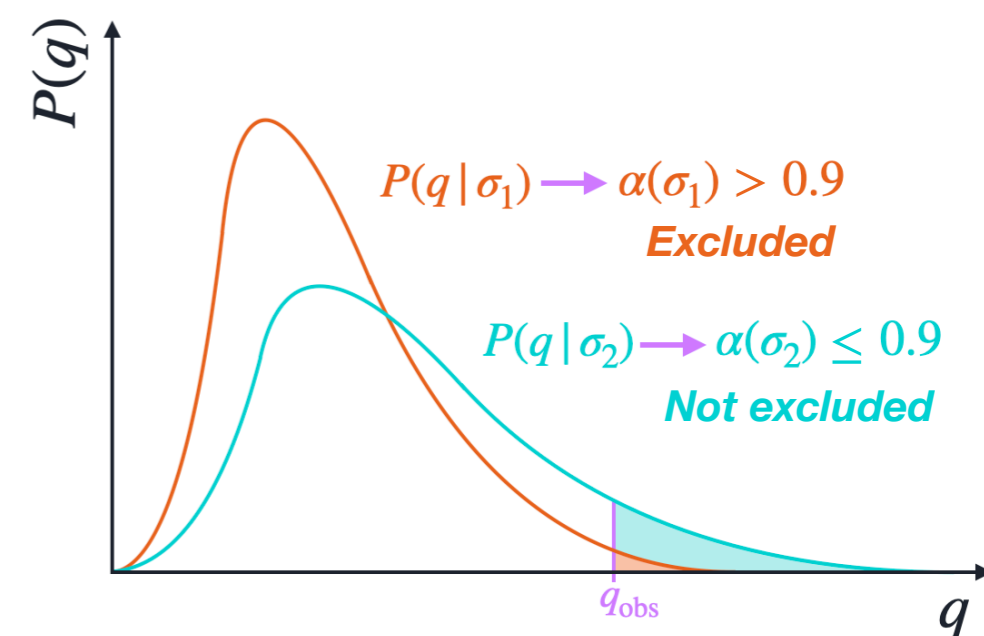
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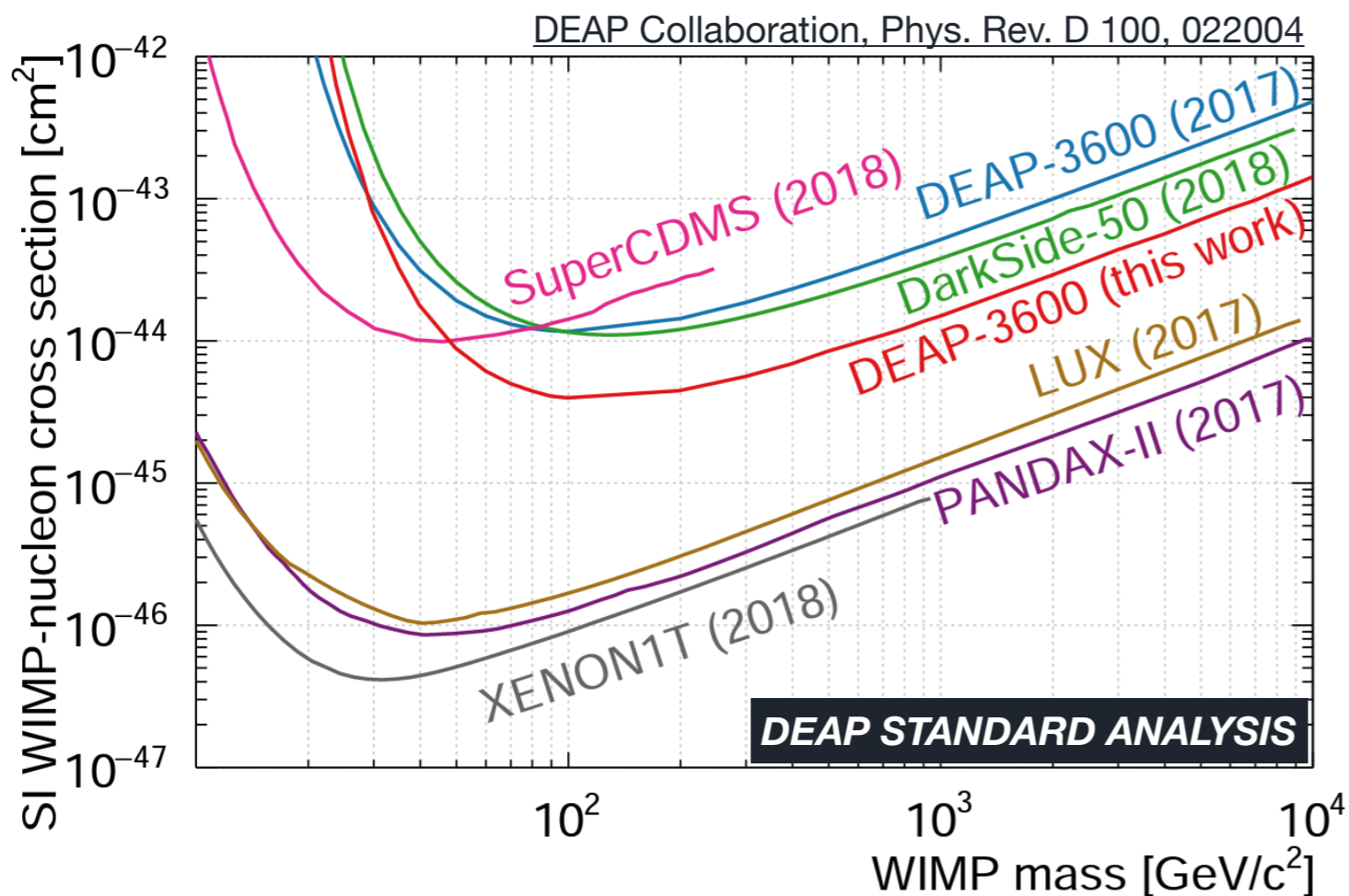
$$\lambda = \frac{\mathcal{L}(\mathbf{x} | \sigma, \hat{\theta})}{\mathcal{L}(\mathbf{x} | \hat{\sigma}, \hat{\theta})} \rightarrow q = \begin{cases} -2\ln\lambda & \sigma \geq \hat{\sigma} \\ 0 & \sigma < \hat{\sigma} \end{cases}$$

Best fit for a given value of  $\sigma$

Best possible fit

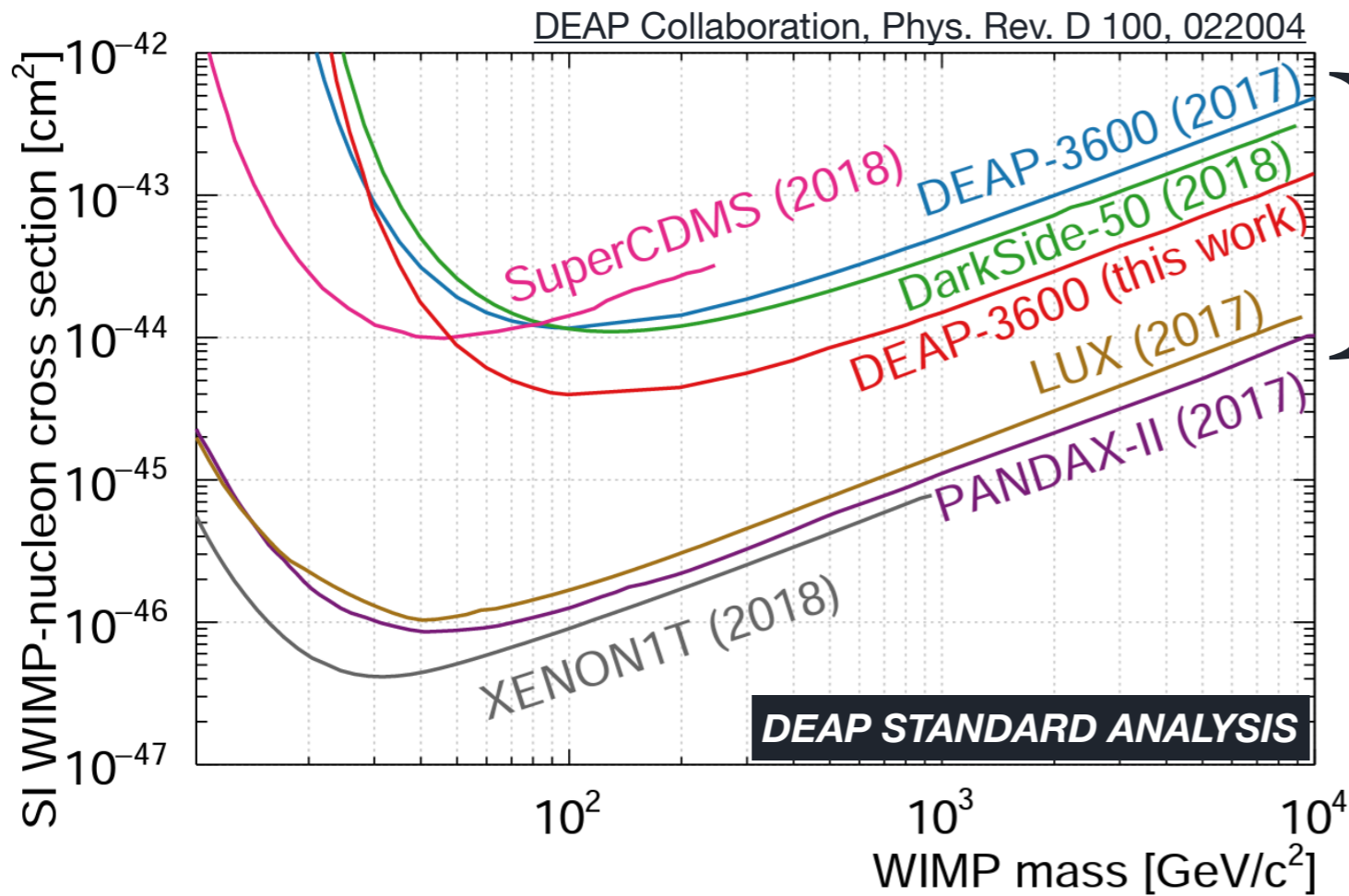


# WIMP Searches: Profile Likelihood Ratio



DEAP sensitivity from 231-day exposure can be improved with PLR and Machine Learning

# WIMP Searches: Nonrelativistic Effective Field Theory



All assuming a scalar WIMP-nucleon coupling; i.e. coherent scattering with entire nucleus

A more general non-relativistic effective field theory includes velocity and spin dependent mechanisms

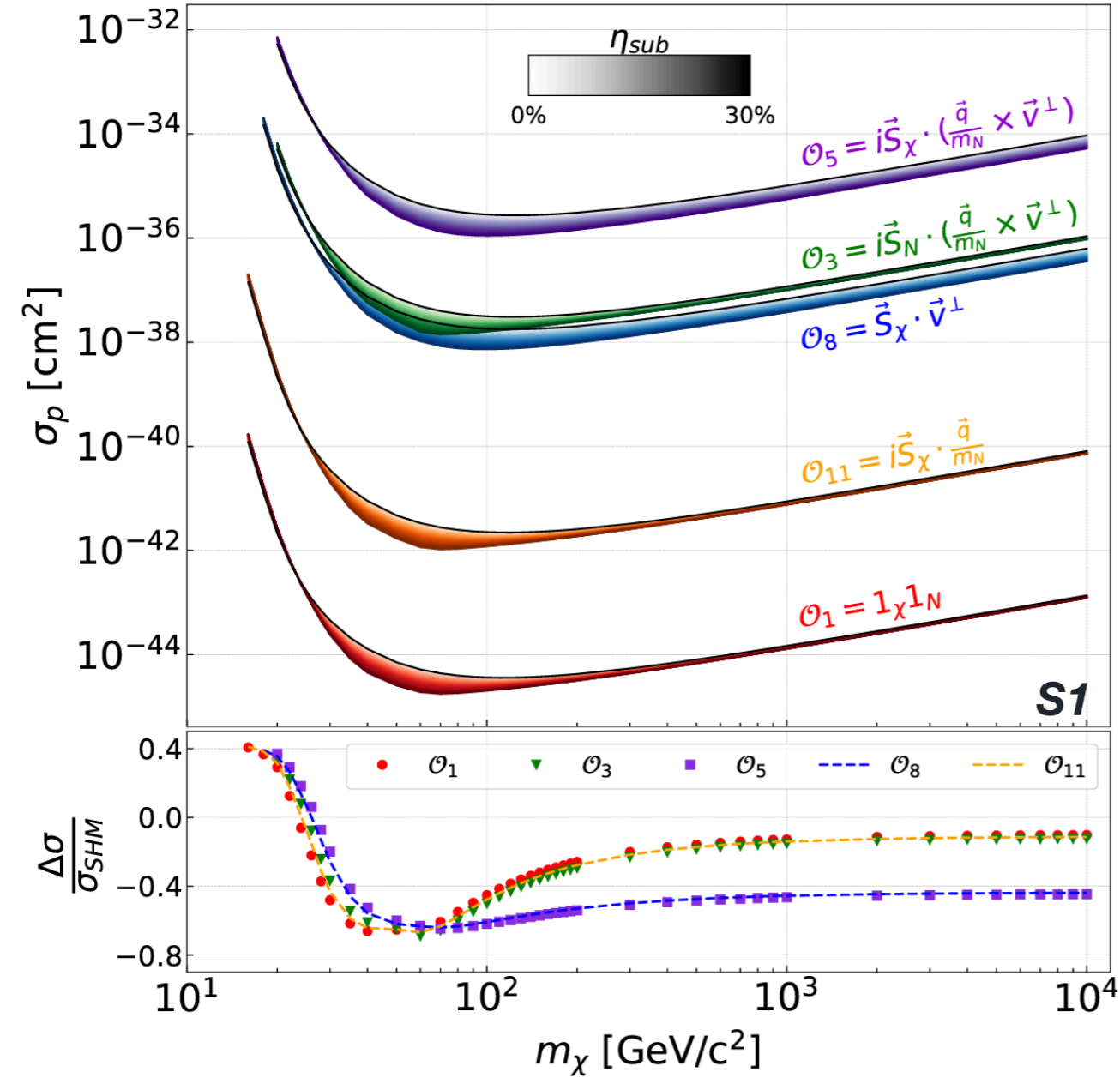
$\mathcal{O}_1$	$1_\chi 1_N$	$\mathcal{O}_{11}$	$iS_\chi \cdot \frac{\vec{q}}{m_N}$
$\mathcal{O}_3$	$iS_N \cdot \left( \frac{\vec{q}}{m_N} \times \vec{v}_\perp \right)$	$\mathcal{O}_{12}$	$\vec{v}_\perp \cdot (S_\chi \times S_N)$
$\mathcal{O}_5$	$iS_\chi \cdot \left( \frac{\vec{q}}{m_N} \times \vec{v}_\perp \right)$	$\mathcal{O}_{15}$	$-\left( S_\chi \cdot \frac{\vec{q}}{m_N} \right) \left[ (S_N \times \vec{v}_\perp) \cdot \frac{\vec{q}}{m_N} \right]$
$\mathcal{O}_8$	$S_\chi \cdot \vec{v}_\perp$		



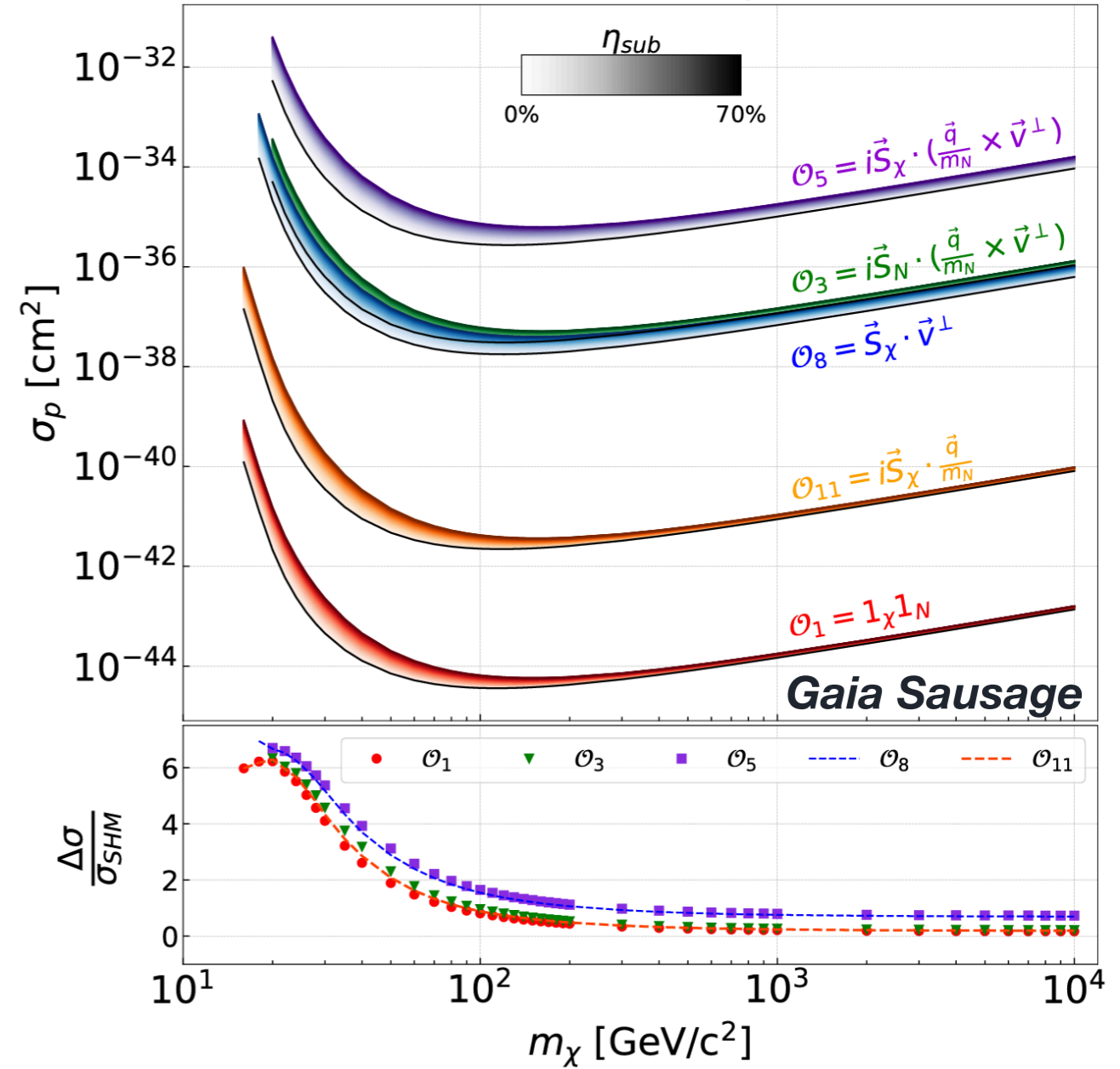
# WIMP Searches: Nonrelativistic Effective Field Theory



DEAP Collaboration Phys. Rev. D 102, 082001



DEAP Collaboration Phys. Rev. D 102, 082001



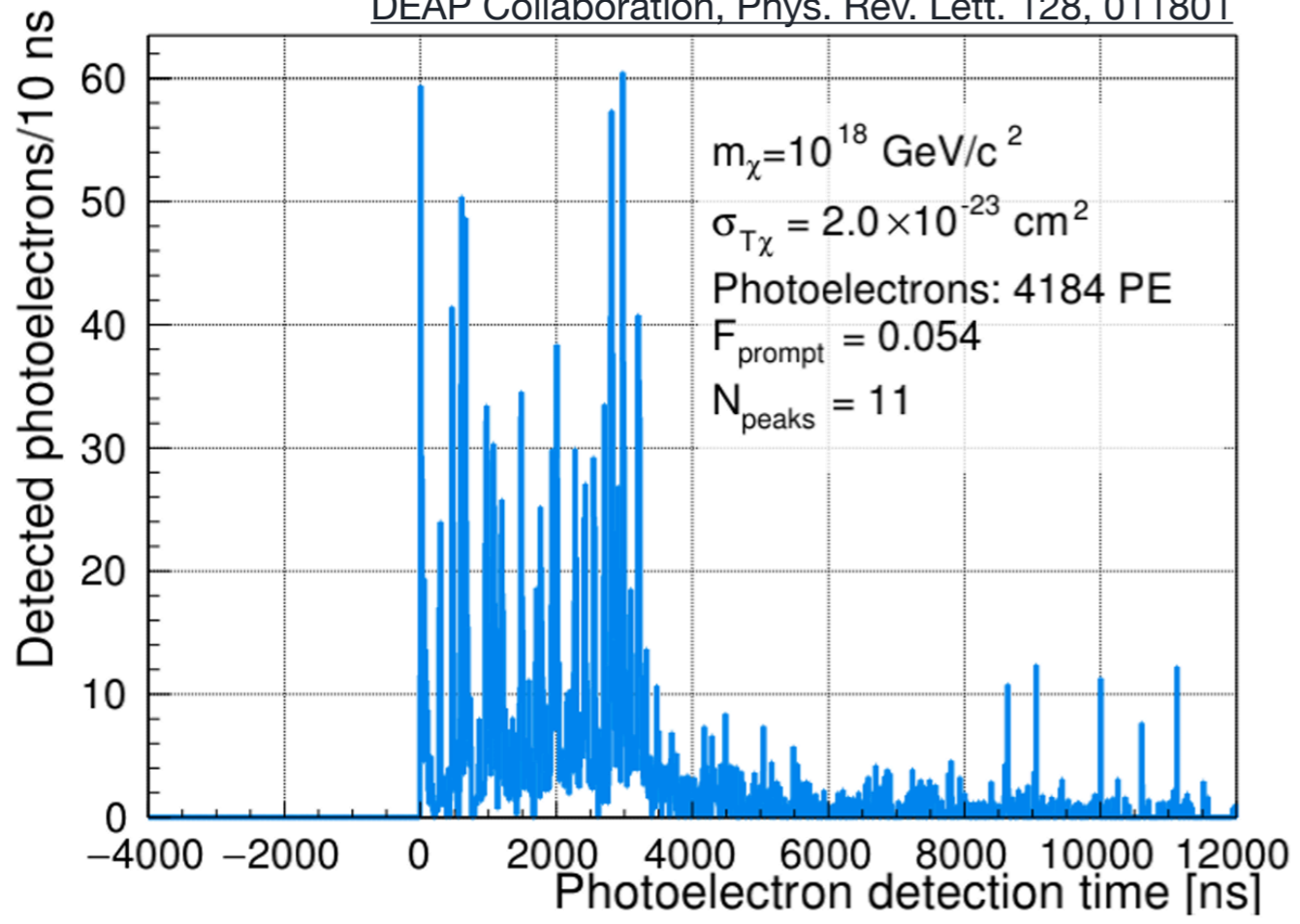
- Interactions in generalized NR-EFT explored with various extensions of standard halo model; substructures like S1 retrograde stellar stream and *Gaia Sausage* considered

# Beyond WIMPs: Planck Scale Dark Matter

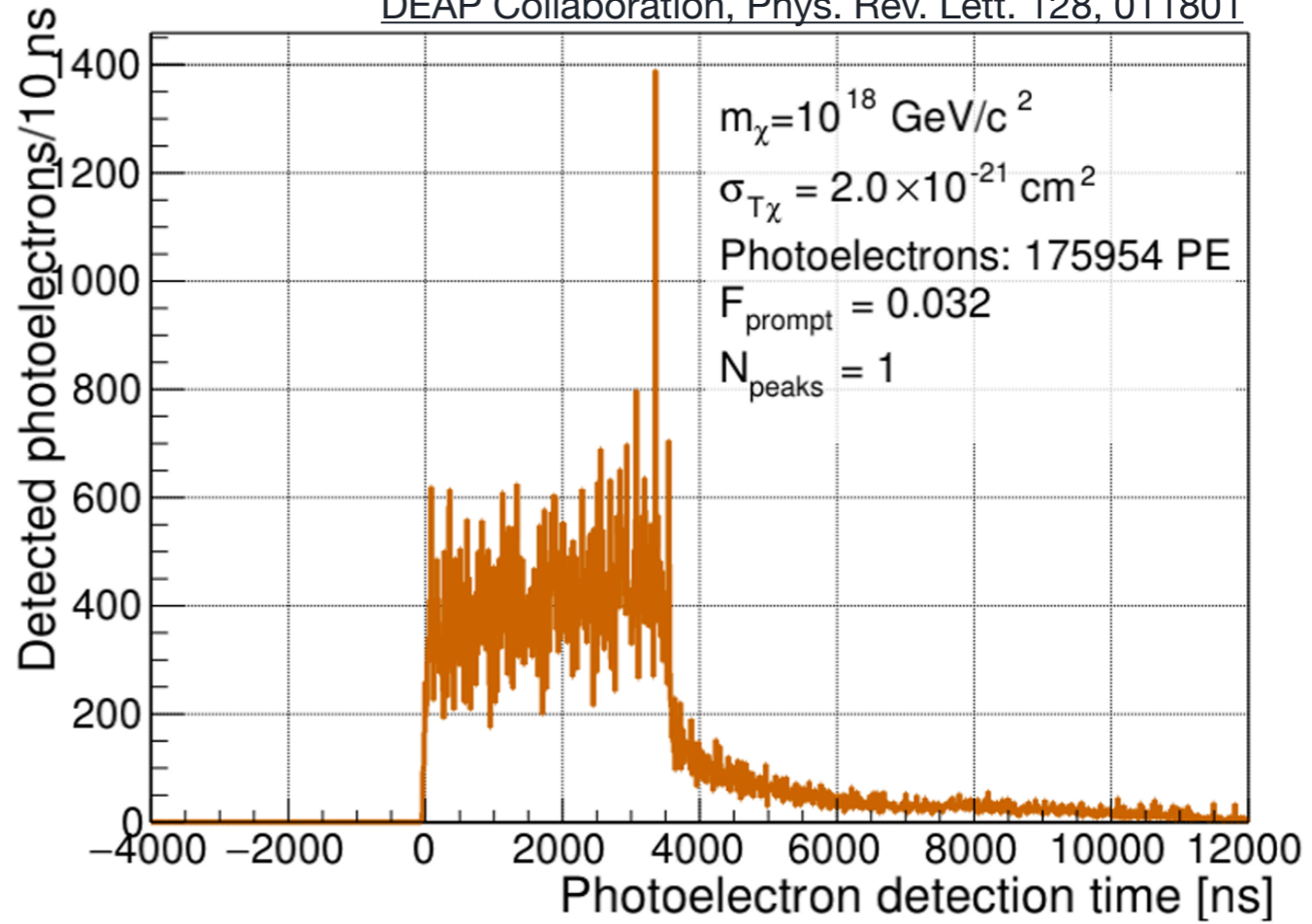


- Dark matter with Planck scale mass is theoretically well motivated; could have much higher cross-sections than WIMPs and not yet be excluded
- Higher cross-sections  $\rightarrow$  multiply scattering DM, which is usually cut in WIMP searches

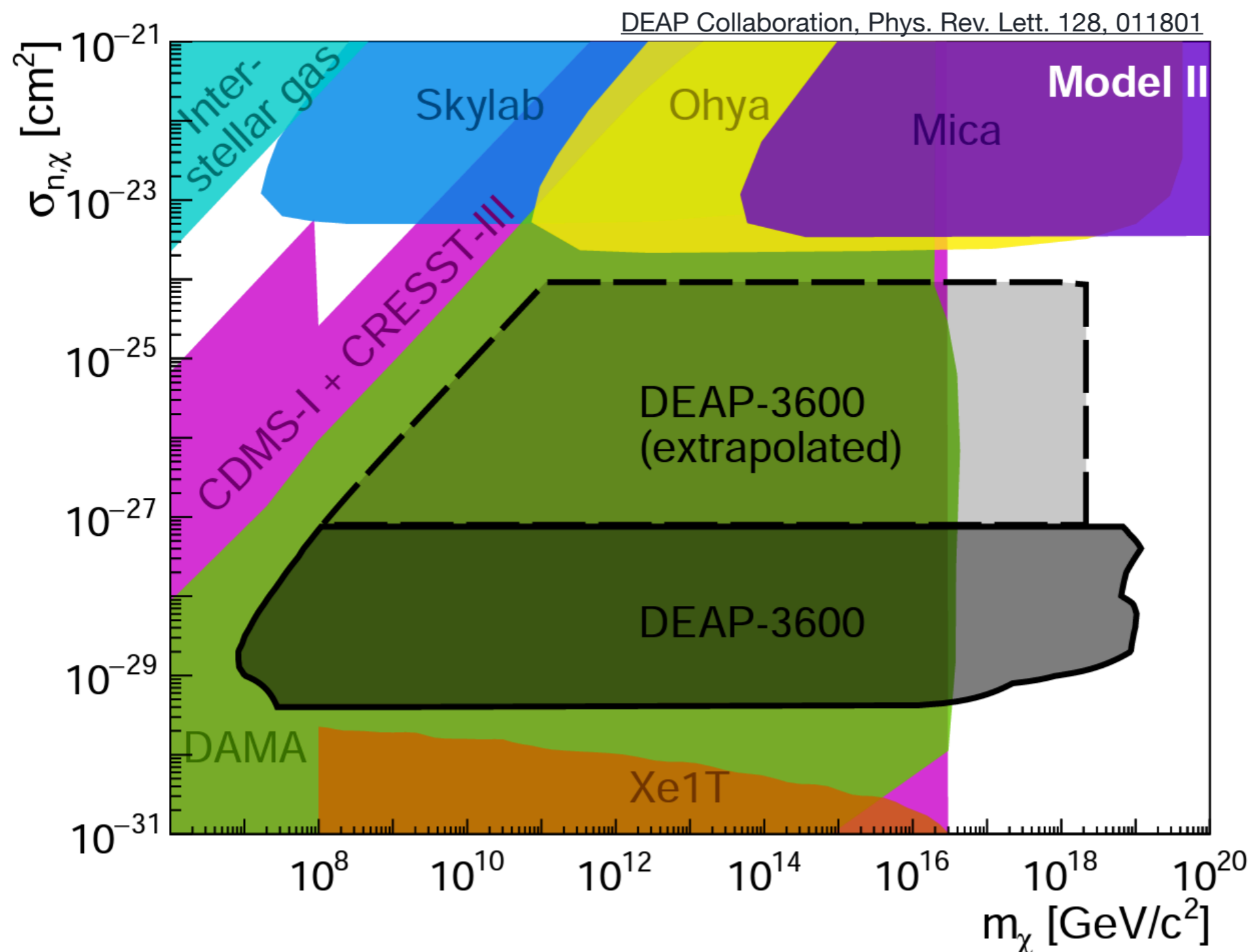
DEAP Collaboration, Phys. Rev. Lett. 128, 011801



DEAP Collaboration, Phys. Rev. Lett. 128, 011801



# Beyond WIMPs: Planck Scale Dark Matter



***First ever direct detection constraint on  
Planck scale dark matter!***

## ***<sup>39</sup>Ar Specific Activity and Half-Life***

- Dedicated papers for <sup>39</sup>Ar specific activity and half-life measurements in DEAP are currently under collaboration review
- Extra slides available for those interested!

## ***5.5 MeV Solar Axion Search***

- Search for axions produced in sun's core via the reaction:  $p + {}^2\text{H} \rightarrow {}^3\text{He} + a$
- Requires precise knowledge of EM backgrounds in MeV range

## ***<sup>8</sup>B Neutrino Absorption***

- DEAP has an active search for inverse beta decay of <sup>40</sup>Ar induced by <sup>8</sup>B solar neutrinos via  $\nu_e + {}^{40}\text{Ar} \rightarrow {}^{40}\text{K}^* + e^-$
- Currently working on background model for this signal, understanding detector response at high energies (4–18 MeV)

## ***Muon Flux at SNOLAB***

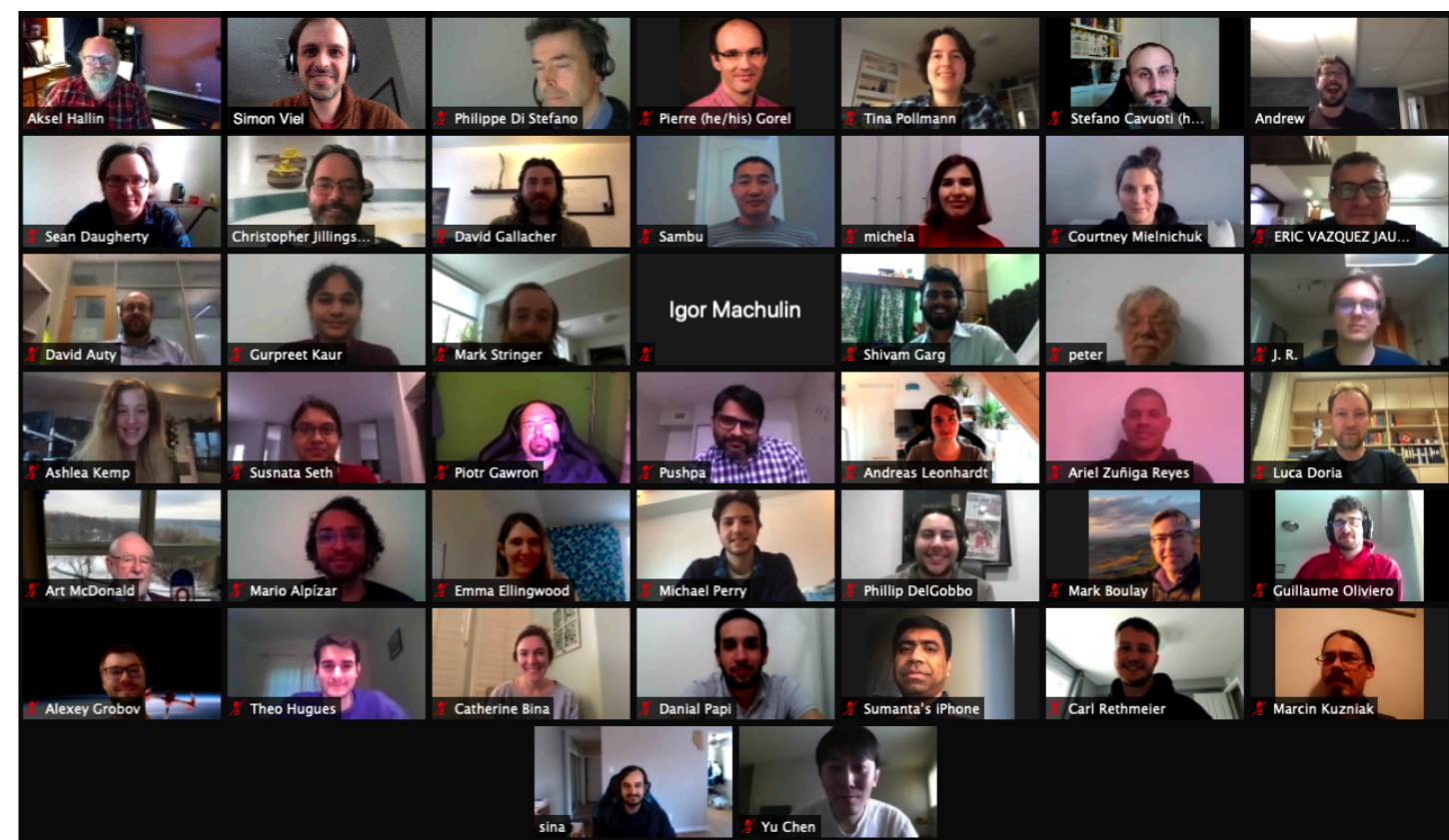
- Dedicated group on DEAP working on muon veto instrumentation paper as well as a muon flux measurement at SNOLAB
- Currently validating MC model, developing event selection criteria to eliminate instrumental events, studying systematics

- Precise LAr pulse shape measurements contribute to excellent background rejection
  - ▶ World leading PSD!
- Competitive dark matter searches spanning 17 orders of magnitude in mass
  - ▶ 100 GeV WIMP search extended with NR-EFT
  - ▶ Previously unprobed Planck Scale DM parameter space excluded at  $10^{19}$  GeV
- Ongoing analyses aimed at improving sensitivity to WIMPs and other new physics
  - ▶ PLR and Machine Learning analyses are well along their way!

# DEAP Collaboration



ROYAL HOLLOWAY UNIVERSITY OF LONDON





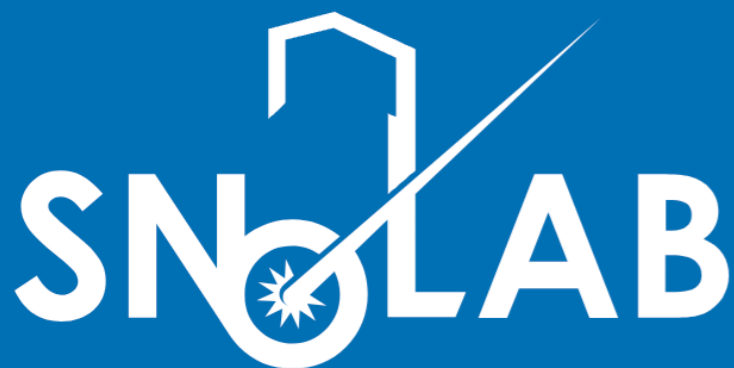
**Grazie per  
l'attenzione!**



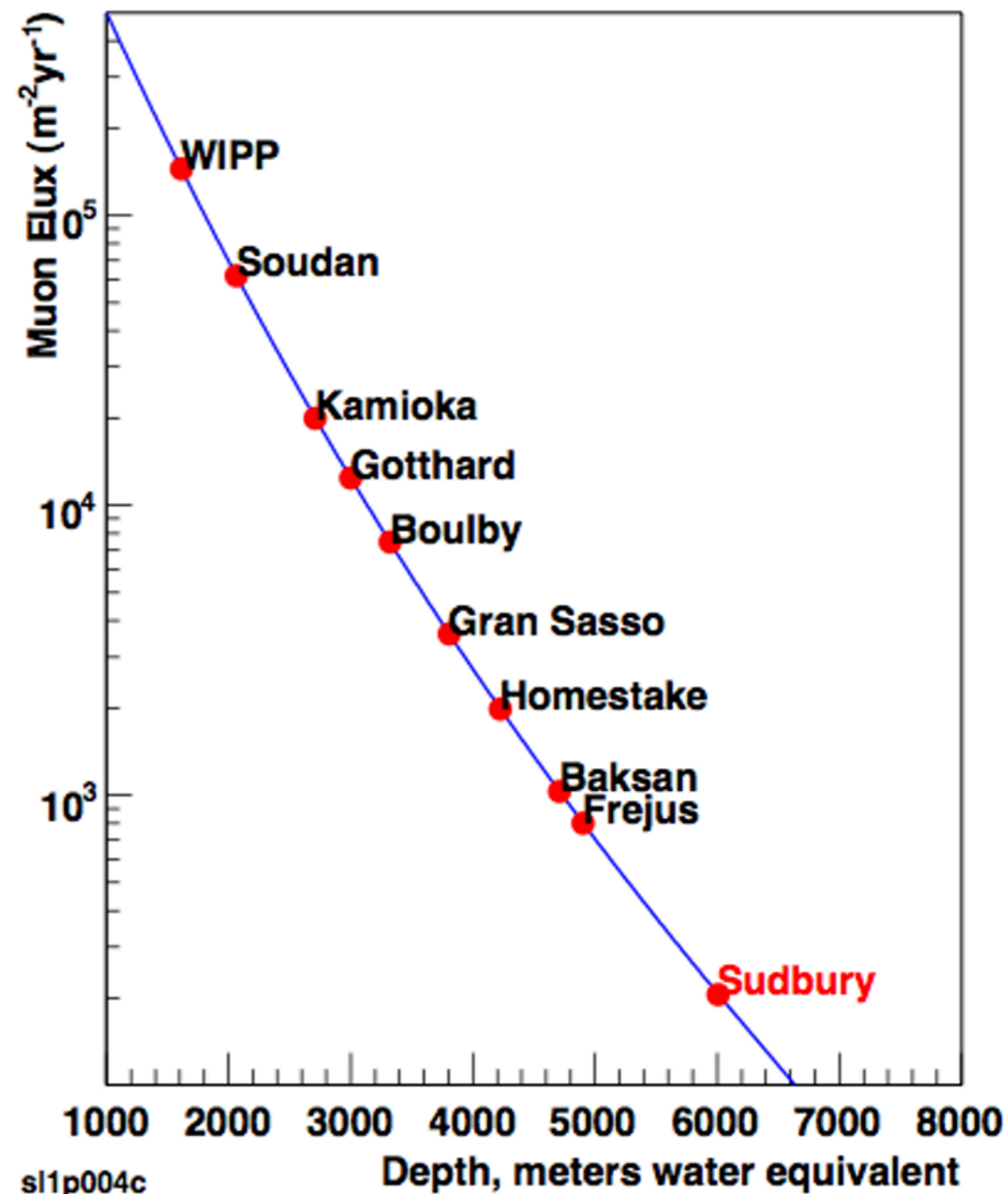
**ROYAL  
HOLLOWAY  
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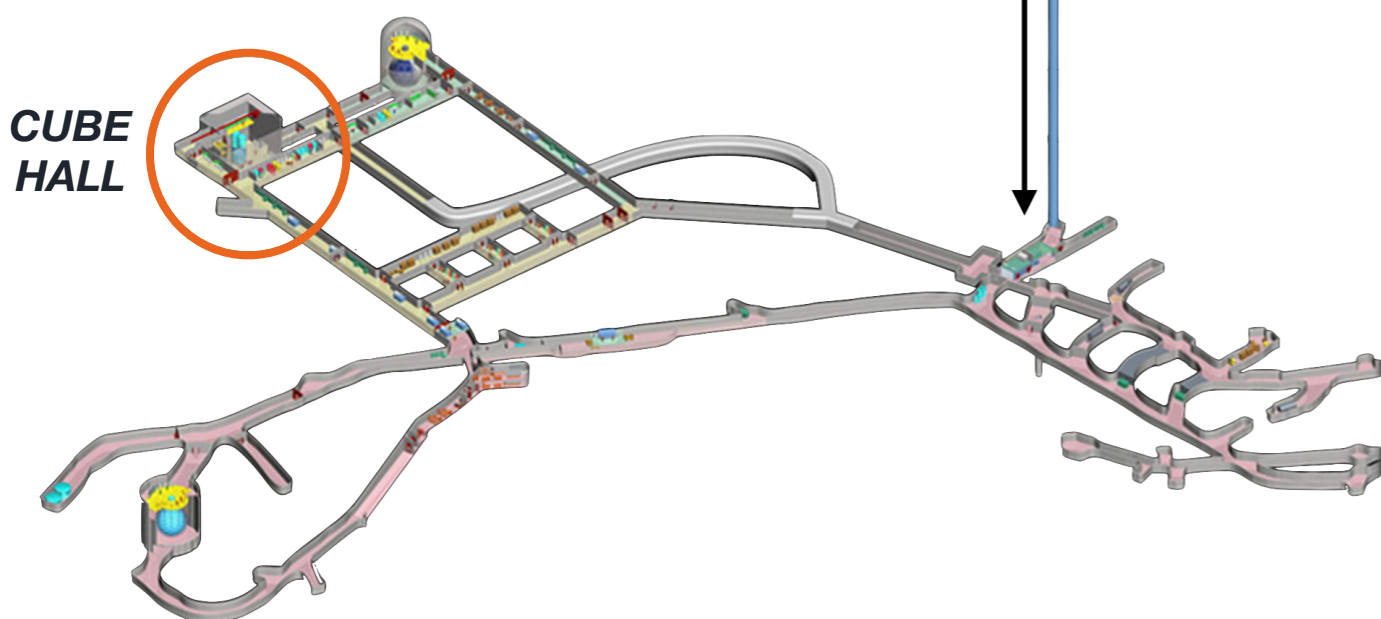
# Extra Slides: SNOLAB



Muon Flux =  $0.27/\text{m}^2/\text{day}$



2 km



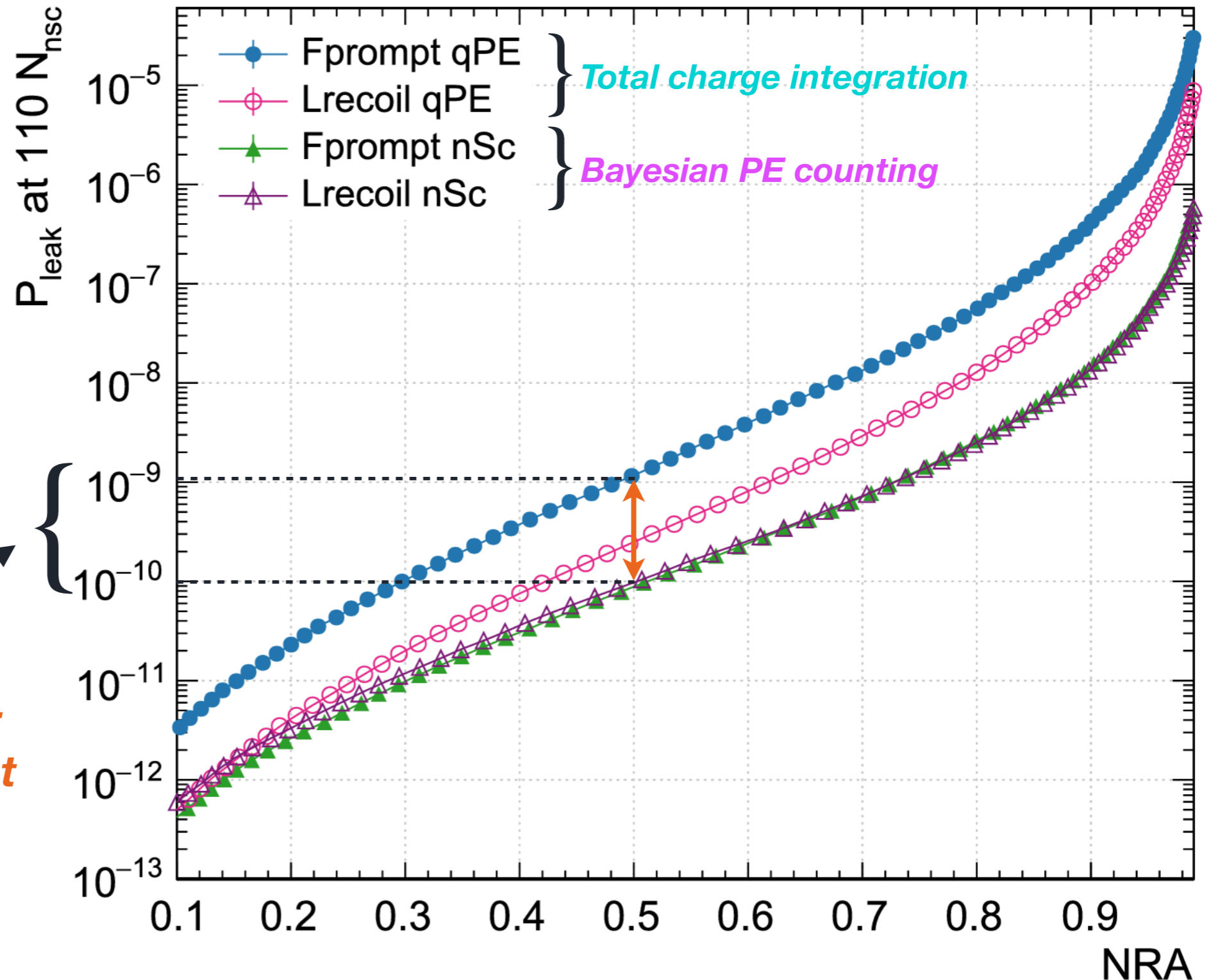


# Extra Slides: <sup>39</sup>Ar Specific Activity and Half-Life



P. Adhikari, et al. Eur. Phys. J. C 81, 823 (2021)

- DEAP has two energy estimators and two PSD models, total of four unique algorithms
- Our best algorithms use the Bayesian PE counting estimator

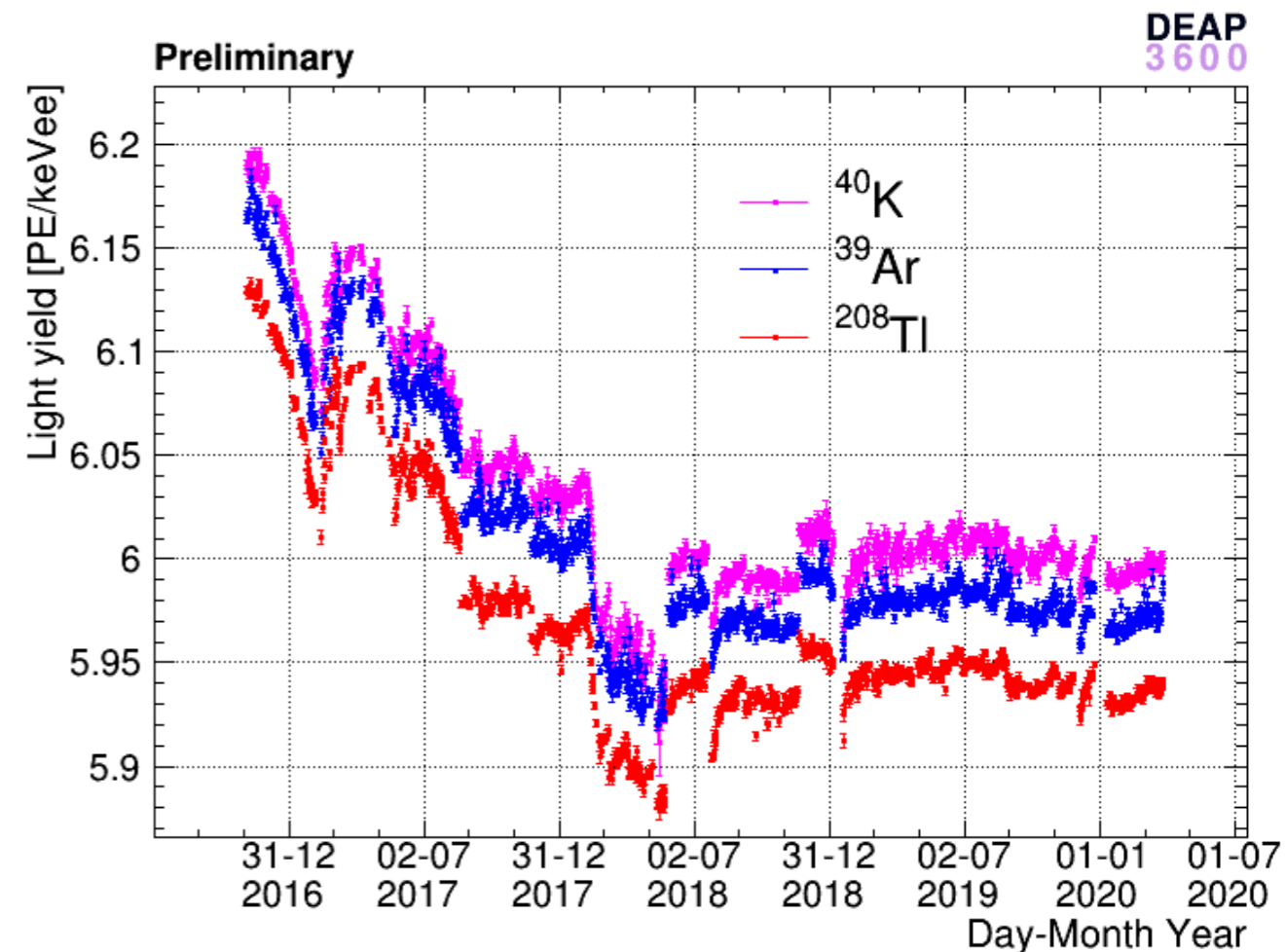
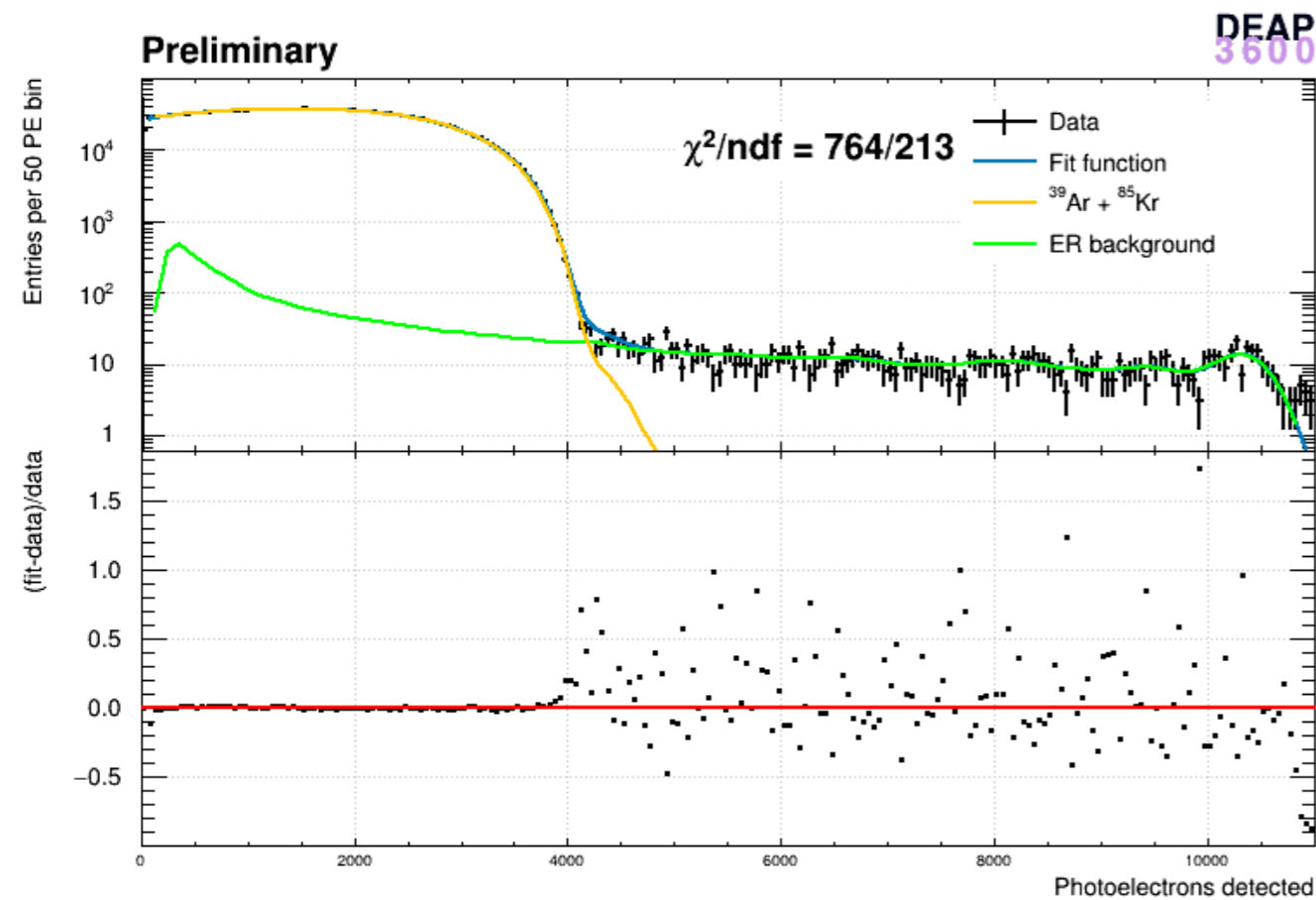


**10x reduction of <sup>39</sup>Ar leakage from simplest algorithm**

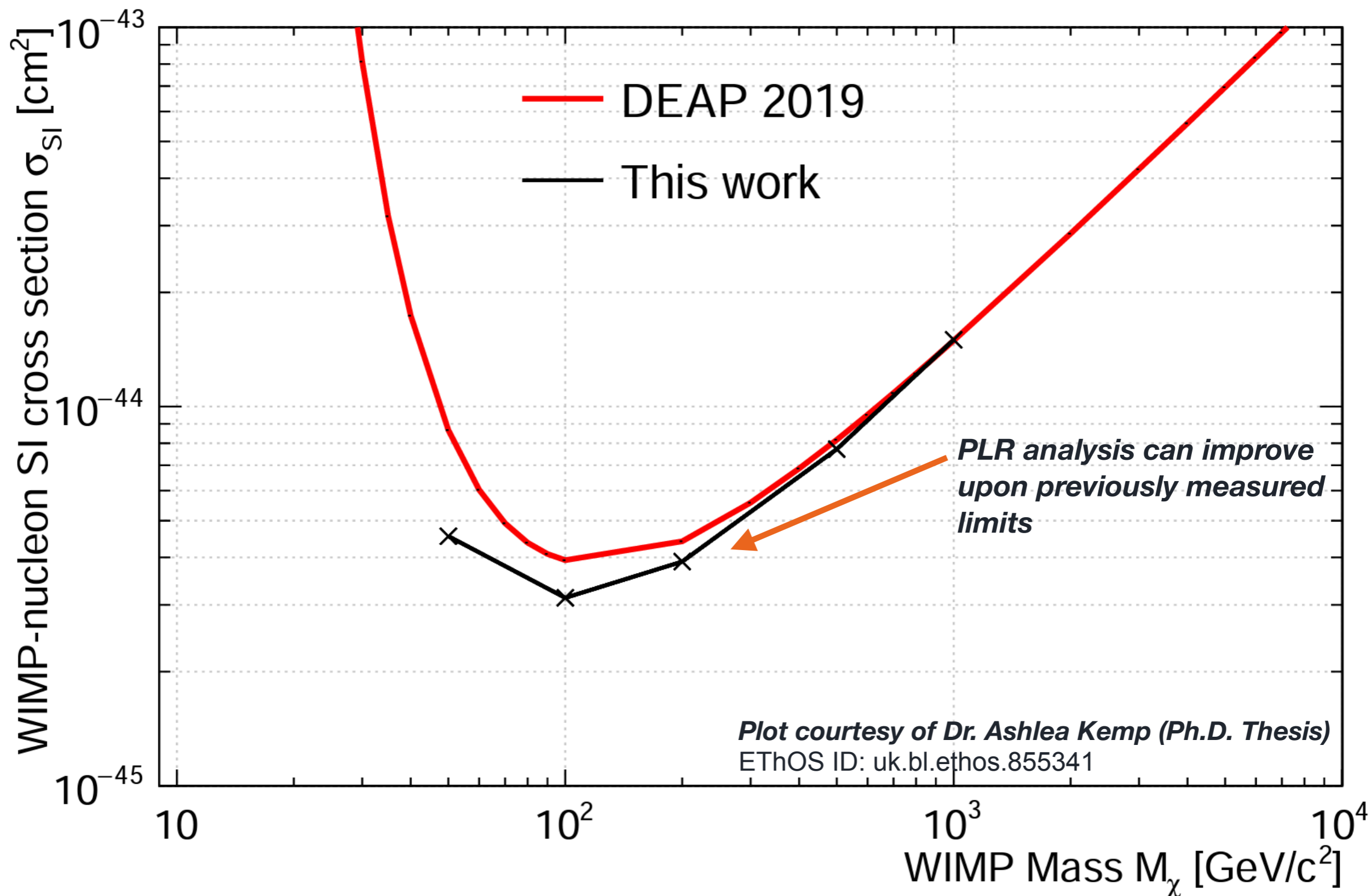
# Extra Slides: <sup>39</sup>Ar Specific Activity and Half-Life



- Dedicated papers for <sup>39</sup>Ar specific activity and half-life in DEAP are currently under collaboration review
- low energy beta spectrum model accounts for <sup>39</sup>Ar and <sup>85</sup>Kr betas, low energy ER band backgrounds, pileup with various other sources
- Drifting of light yield also included in systematic analysis; stable to within ~0.3 PE/keV<sub>ee</sub>



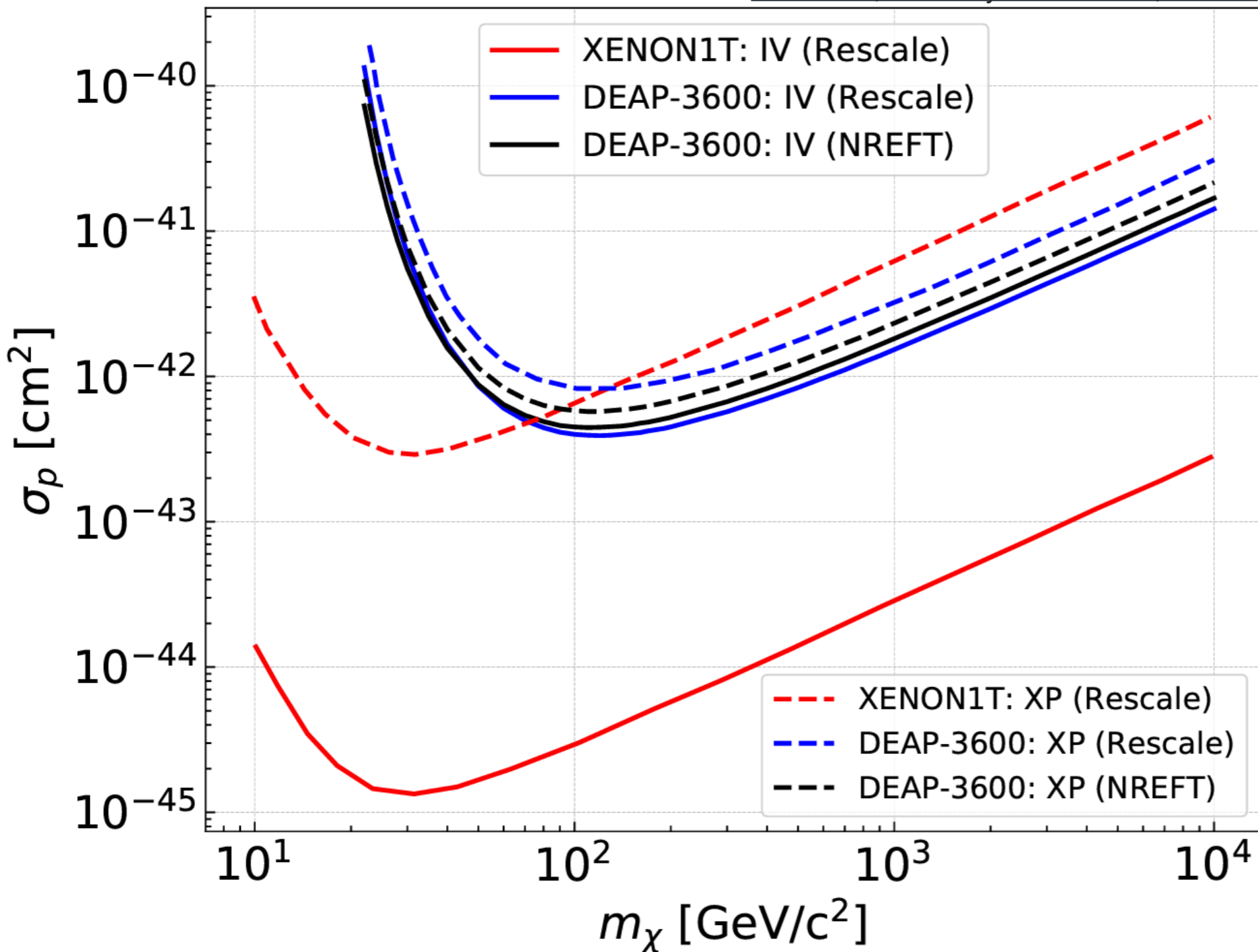
# Extra Slides: Profile Likelihood Ratio



# Extra Slides: Xenophobic WIMPs



P. Adhikari, et al. Phys. Rev. D 102, 082001



- Isospin-violating interactions also considered in NR-EFT framework
- *xenophobic* (XP) interactions cover a range of isospin-violating models
- DEAP sets world leading limit on these isospin-violating interactions

# Extra Slides: Planck Scale Dark Matter

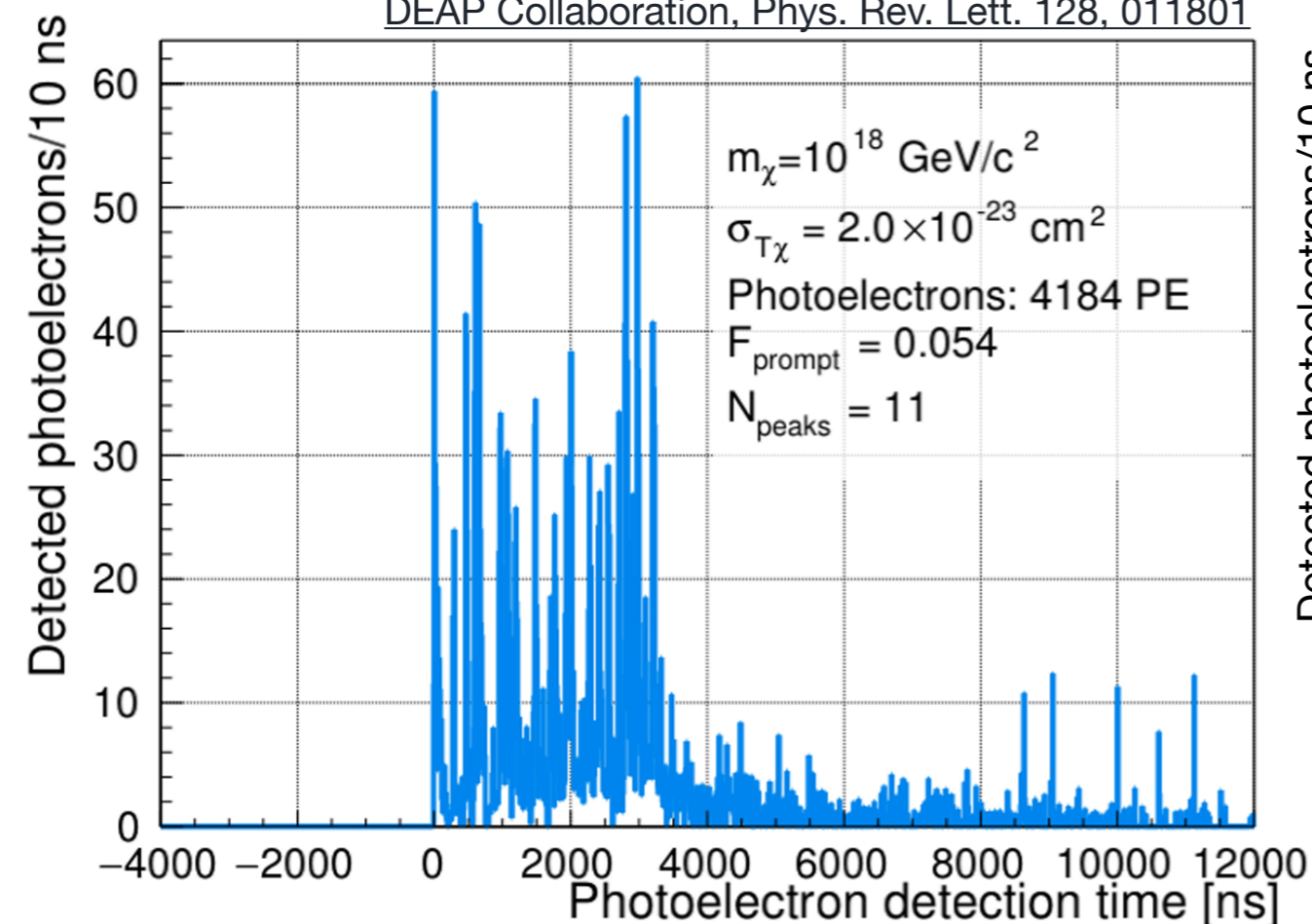


- Dark matter with Planck scale mass is theoretically well motivated; could have much higher cross-sections than WIMPs and not yet be excluded
- Higher cross-sections  $\rightarrow$  multiply scattering DM, which is usually cut in WIMP searches

***Distinguishable from pileup***

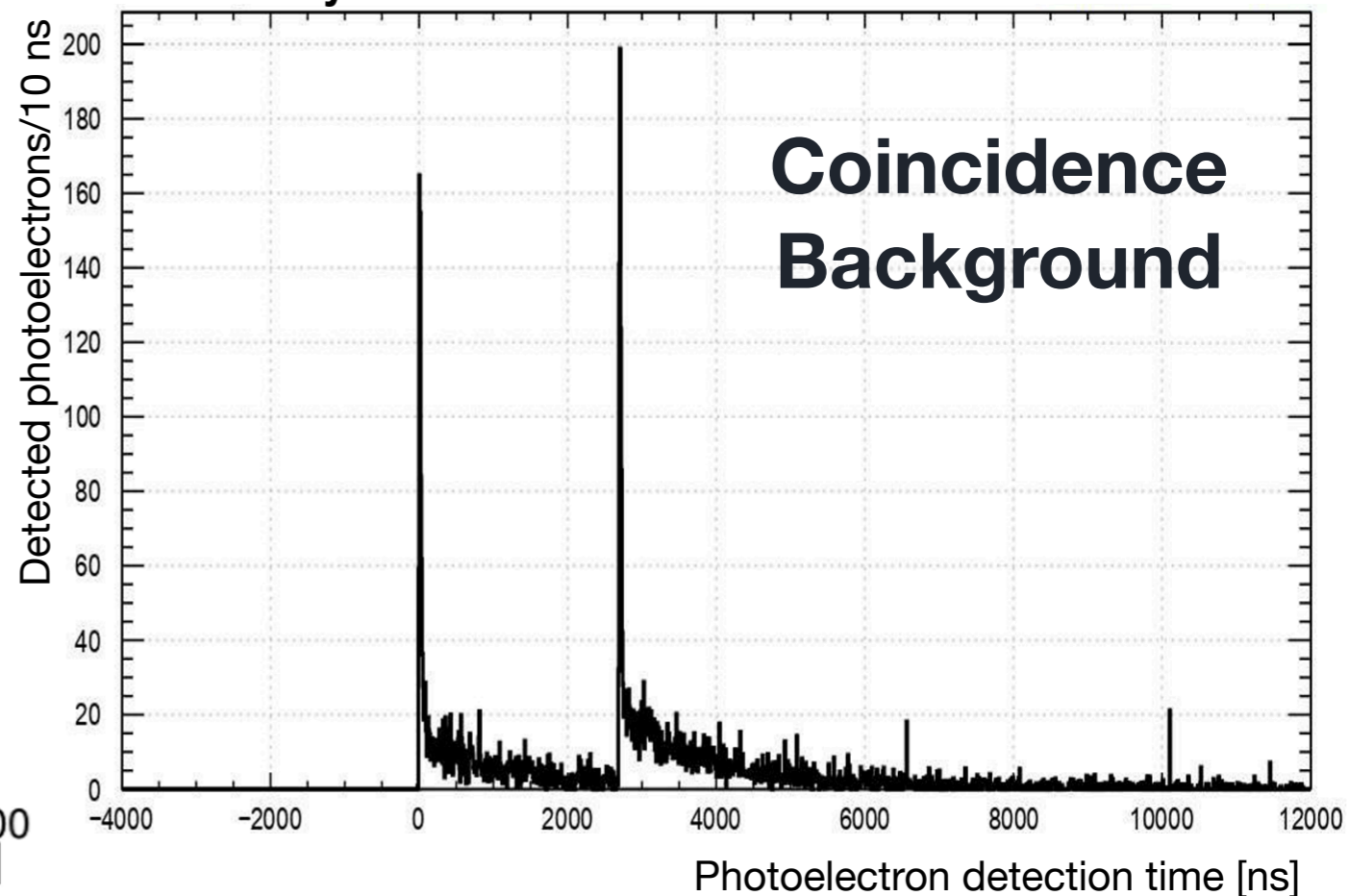


DEAP Collaboration, Phys. Rev. Lett. 128, 011801



Preliminary

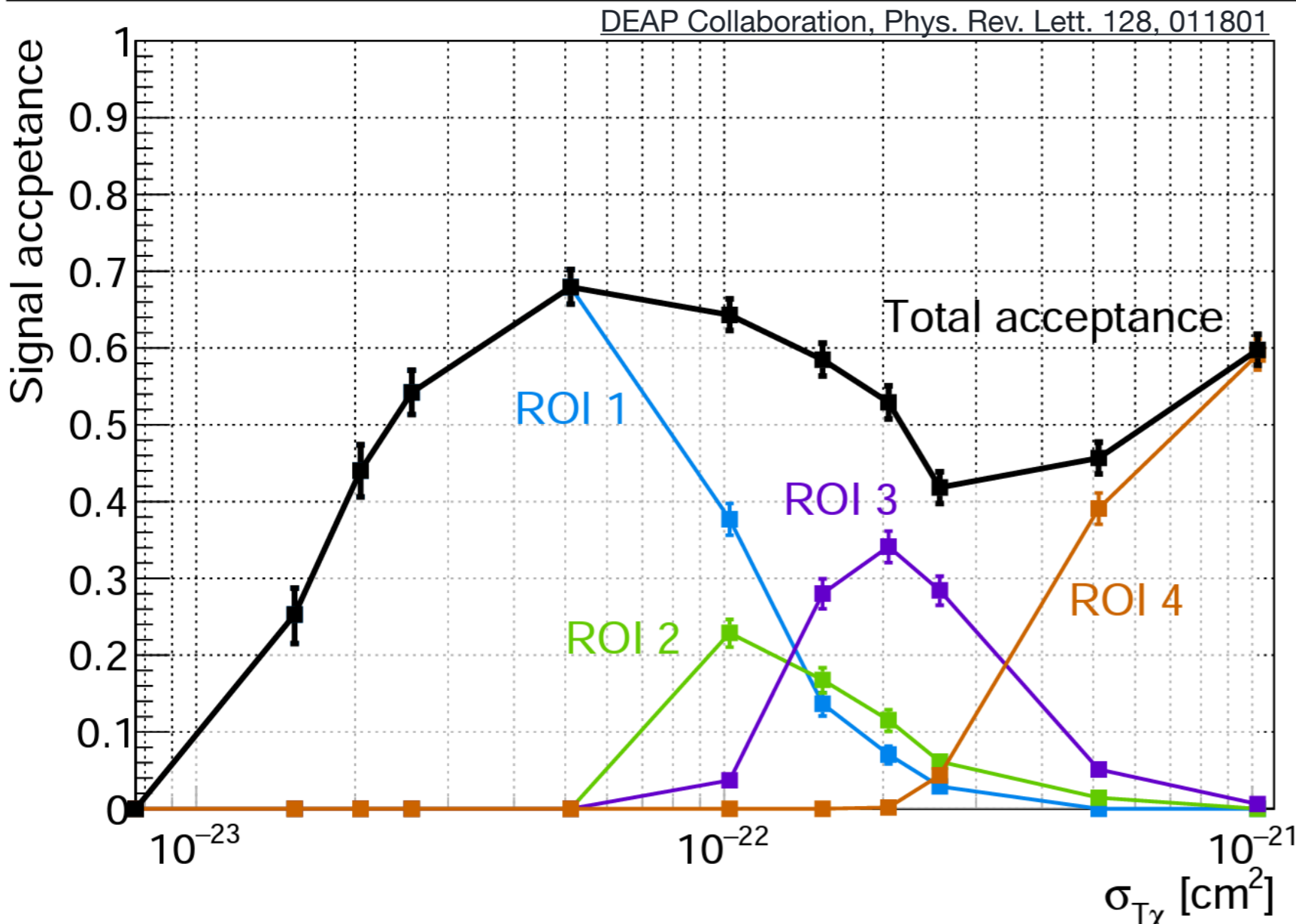
DEAP  
Simulation 3600



# Extra Slides: Planck Scale Dark Matter



ROI	PE range	Energy [MeV <sub>ee</sub> ]	N <sub>peaks</sub> <sup>min</sup>	F <sub>prompt</sub> <sup>max</sup>	$\mu_b$	N <sub>obs.</sub>
1	4000–20 000	0.5–2.9	7	0.10	$(4 \pm 3) \times 10^{-2}$	0
2	20 000–30 000	2.9–4.4	5	0.10	$(6 \pm 1) \times 10^{-4}$	0
3	30 000–70 000	4.4–10.4	4	0.10	$(6 \pm 2) \times 10^{-4}$	0
4	70 000– $4 \times 10^8$	10.4–60 000	0	0.05	$(10 \pm 3) \times 10^{-3}$	0



- Look for events with multiple peaks and/or very low  $F_{\text{prompt}}$
- Defined 4 ROIs with high signal acceptance
- Backgrounds in these ROIs are negligible;  $\ll 1$  event in 813 live days

# Extra Slides: Planck Scale Dark Matter



- Model I considers the case where:  $\frac{d\sigma_{T\chi}}{dE_R} = \frac{d\sigma_{n\chi}}{dE_R} |F_T(q)|^2$
- Model II considers the case where:  $\frac{d\sigma_{T\chi}}{dE_R} \approx \frac{d\sigma_{n\chi}}{dE_R} A^4 |F_T(q)|^2$
- DEAP sets new world leading constraints for Planck Scale DM in both scenarios

