

## DEAP-3600 Dark Matter Search at SNOLAB: Overview, Status and Future

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# Liquid argon as a robust and scalable dark matter target



- Well-separated singlet and triplet lifetimes in argon allow for good pulse-shape discrimination (PSD) of β/γ's using only scintillation time information
- PSD to 10<sup>-8</sup> demonstrated with DEAP-1 (Astroparticle Physics 25, 179 (2006) and arXiv:0904.2930, analysis of extended dataset to be published)
- For DEAP-3600 projected to 10<sup>-10</sup> at 15 keVee
- Very large target masses possible, since no absorption of UV scintillation photons in argon, and no e-drift requirements.
- 1000 kg argon target allows 10<sup>-46</sup> cm<sup>2</sup> sensitivity (SI) with ~15 keVee (60 keVr) threshold, 3-year run





## Background targets



Background	Target
Radon in argon	< 1.4 nBq/kg
Surface $\alpha$ 's (tolerance using conservative pos. resolution)	< 0.2 μBq/m2
Surface $\alpha$ 's (tolerance using ML position resolution)	< 100 µBq/m2
Neutrons (all sources, in fiducial volume)	< 2 pBq/kg
Bg events, dominated by 39Ar	< 2 pBq/kg
Total Backgrounds (3 Tonne-year in fiducial volume and Region of Interest)	< 0.6 events



#### Radon backgrounds in the DEAP-1 liquid argon based Dark Matter detector

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#### arXiv:1211.0909v1 [astro-ph.IM] 5 Nov 2012





Low-energy spectrum well-described by  $^{222}\text{Rn}$  in argon, normalized to high-energy  $\alpha$ -rates.

Gap between 15 keVee and 40 keVee for DEAP-3600.

#### By-product:

"Surface roughness interpretation of CRESST-II result"

Astropart. Phys. 36, 77 (2012)

# In DEAP-3600 surface background better discriminated with fiducialization



## DEAP-3600 neutron backgrounds

- Dominated by (alpha, n) neutrons from PMT glass (Hamamatsu R5912 HQE)
- Extensive Geant4 simulations to set the purity targets for U/Th and <sup>210</sup>Pb
- Quality assurance and assay campaign to validate the material purity and limit exposure to Rn

	# of neutrons (produced in 3 years)	Events in ROI (3 years)
Acrylic vessel	<44 (Ge g-assay)	<0.096
Light guides	<127 (Ge g-assay)	<0.015
Filler blocks	<173 (Ge g-assay)	<0.034
PMTs	2.6x105	0.140
PMT mounts	7565	0. 010
Rn emanation	<44	<0.081
Rn deposition (3 months construction)	38	0.010
Other sources		0.04
Total	<2.7x105	<0.35





### **DEAP-3600**



3600 kg argon target (1000 kg fiducial) in sealed ultraclean Acrylic Vessel

Vessel is "resurfaced" in-situ to remove deposited Rn daughters after construction

255 Hamamatsu R5912 HQE PMTs 8-inch (32% QE, 75% coverage)

50 cm light guides + PE shielding provide neutron moderation

Detector in 8 m water shield at SNOLAB

9









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м. кизпак (Snowmass Cosmic Frontier Workshop)

## Construction highlights





### Water shield tank











(Snowmass







AV Fabrication (RPT and University of Alberta)







# Underground bonding and machining







## Cooldown scheduled for January 2014



- Remaining construction milestones:
  - LG bonding
  - PMT installation
  - Resurfacing
  - TPB deposition
- Cooldown: January 2014
- Followed by a couple of months for commissioning

## Thinking about future scale-up



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## WIMP mass sensitivity





- Technology can be scaled to very large target masses, > 100 tonnes or 10<sup>-48</sup> cm<sup>2</sup> sensitivity
- Larger detector allows for better position reconstruction
- This makes surface contamination easier to mitigate
- Relaxed targets on surface contamination significantly simplify many aspects of construction and assembly (compared to DEAP-3600)
- Large detector will require Depleted
   Argon

## Summary



- DEAP-3600 construction is progressing rapidly
- Detector online early next year, with competitive sensitivity for WIMP masses >150 GeV
- We have demonstrated sufficient control over surface backgrounds in DEAP-1

- Some conceptual effort on the next generation detector
- In the single-phase technology, larger scale makes life much easier
- Potentially, very attractive way towards a precision measurement (if a WIMP signal is seen by 1 tonne scale experiments)